

Donauer, Stefanie; Heinen, Florian; Sibbertsen, Philipp

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Identification problems in ESTAR models and a new model

Stefanie Donauer, Florian Heinen and Philipp Sibbertsen*

Discussion Paper No. 444

Leibniz Universität Hannover
Institute of Statistics
Königsworther Platz 1, 30167 Hannover, Germany

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Abstract

In ESTAR models it is usually difficult to determine parameter estimates, as it can be observed in the literature. We show that the phenomena of getting strongly biased estimators is a consequence of the so-called identification problem, the problem of properly distinguishing the transition function in relation to extreme parameter combinations. This happens in particular for either very small or very large values of the error term variance. Furthermore, we introduce a new alternative model -the T-STAR model- which has similar properties as the ESTAR model but reduces the effects of the identification problem. We also derive a linearity and a unit root test for this model.

JEL-Numbers: C12, C22, C52

Keywords: Nonlinearities · Smooth transition · Linearity testing · Unit root testing · Real exchange rates

1 Introduction

Nonlinear time series models have become more and more popular over the last decade. In particular, Exponential Smooth Transition Autoregressive (ESTAR) models have been used for modeling real exchange rates. These models contain of two autoregressive regimes which are connected by a smooth transition function of an exponential type. Under certain regularity conditions they are globally stationary. This is even the case if one regime is assumed to be a random walk as it happens for real exchange rates where we assume one regime to have a unit root whereas the other regime is a stationary autoregressive process. Moreover, the U-shape of the transition function is a desired property in the context of real exchange rates as there one wants to allow the exchange rates to move freely like a random walk near an equilibrium and being pulled back to it once they move too far away from it.

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Contrary to linear models where parameter estimates are independent of the size of the error variable, we face an identification problem in ESTAR models due to the influence of the error term variance. Small or large error variances no longer allow to identify the ESTAR model in the sense that they stay only in one of the two regimes and do not switch between the regimes any more. This problem was first observed in [Luukkonen et al. \(1988\)](#) who mention it in a short remark. However, to the best of our knowledge it has not been further considered in the literature since.

In the literature people encounter problems when determining parameter estimates. Various subjective tricks have been proposed (see e.g. [Haggan and Ozaki \(1981\)](#), [Lütkepohl and Krätzig \(2004\)](#)) in order to circumvent these in order to guarantee a better performance of the estimators. We now show that the identification problems causes some parameters to be unidentifiable. In other words, no general estimation procedure will produce reasonable estimators, making some modification to the optimization procedure necessary. This has of course its limits as there is no theory saying that these methods work in general.

The identification problem is also visible in examining linearity tests. Linearity tests against ESTAR have been developed by for example [Teräsvirta \(1984\)](#) and unit root tests against an ESTAR alternative can be found in [Kapetanios et al. \(2003\)](#). Both of these tests have the nonintuitive property of a low power when the error term variance is either very small or very large. For the linearity test this was stated in [Luukkonen et al. \(1988\)](#) and for the unit root test see [Kruse et al. \(2008\)](#). This effect is less surprising for a large error term variance as in this case the noise dominates the signal. However, it is rather surprising in the opposite case of a small error variance as in this case the signal dominates the noise. Therefore, an increase of the power would be expected. As real exchange rates have extremely small error variances (see for example [Taylor et al., 2001](#)) this problem is of a high practical relevance and can lead to false non-rejections of the null and therefore rejecting a nonlinear adjustment process for real exchange rates.

We now introduce a natural alternative of the ESTAR model by using a different transition function, leading to the T-STAR model. This transition function possess the same desired properties and can therefore be applied to the same situations. The new transition function has however fatter tails which turns out to reduce the identification problem. We can improve the estimation procedure for extreme error term variances. In particular, standard optimization tools can be used. Moreover, we develop a linearity and a unit root test for this new model and study their performances in extensive simulations.

The rest of the paper is organized as follows. In the next section we define ESTAR models in more detail and analyze the identification problem, in particular with respect to small error term variances. The new T-STAR model is examined in [Section 3](#). After describing the model (see [Section 3.1](#)) we derive the linearity as well as the unit root test in [Sections 3.2](#) and [3.3](#), respectively. The simulation studies we performed are summarized in [Section 3.4](#). A comparison of the ESTAR and T-STAR model is presented in [Section 4](#), discussing real exchange data. [Section 5](#) concludes whereas all proofs are collected in the [Appendix](#), together with certain technical lemmas.

2 Exponential Smooth Transition Autoregressive Models

In this section we introduce the general Smooth Transition Autogressive (STAR) model. The transition function for ESTAR models is specified and basic properties of the resulting model are studied. Subsequently, the identification problem present in the ESTAR setting is described and analyzed in [Section 2.2](#).

One speaks of a Smooth Transition Autoregressive (STAR) model, if two autoregressive regimes are connected by a transition function satisfying certain smoothness conditions. Here, smoothness is meant that the transition function changes smoothly from zero to one and therefore governs the transition between the two regimes in a smooth way. Alternatively, a STAR model can also be interpreted as a continuum of regimes which is passed through by the process. In general, univariate STAR(p) models, $p \geq 1$, are given by

$$y_t = [\Psi w_t] \times [1 - G(y_{t-d}, \gamma, c)] + [\Theta w_t] \times G(y_{t-d}; \gamma, c) + \varepsilon_t \quad (1)$$

$$= [\Psi w_t] + [\Phi w_t] \times G(y_{t-d}; \gamma, c) + \varepsilon_t, \quad t \geq 1, \quad (2)$$

with initial variable y_0 . The parameter vectors Ψ and Θ as well as w_t are given by $\Psi = (\psi_0, \psi_1, \dots, \psi_p)$, $\Theta = (\vartheta_0, \vartheta_1, \dots, \vartheta_p)$, and $w_t = (1, y_{t-1}, \dots, y_{t-p})'$. For the alternative parametrization (2) we have $\Phi = (\varphi_0, \varphi_1, \dots, \varphi_p) = (\psi_0 - \vartheta_0, \psi_1 - \vartheta_1, \dots, \psi_p - \vartheta_p)$. Different choices of the so-called transition function $G(\cdot; \gamma, c) : \mathbb{R} \rightarrow [0, 1]$ lead to different STAR models. For the rest of this text we assume that the random error terms ε_t satisfy the following conditions:

Assumption 2.1.

The innovations ε_t are assumed to

- (i) *be iid random variables with mean zero and unknown variance σ^2 ,*
- (ii) *have a symmetric density around zero.*

A recent overview of STAR models, estimation techniques and model building procedures can be found in [Franses and van Dijk \(2000\)](#).

Common choices for the function G are the exponential function, leading to the Exponential STAR (ESTAR) model, or the logistic function, depending on the nature of the studied transition. However, the parameter γ is always the transition parameter that governs the speed of the regime changes. For the rest of this document we only consider situations where G is symmetrically U-shaped around the location parameter $c \in \mathbb{R}$ with

$$\lim_{\gamma \rightarrow +\infty} G(\cdot; \gamma, c) \equiv 1 - \mathbf{1}_c, \quad \lim_{\gamma \rightarrow 0} G(\cdot; \gamma, c) \equiv 0 \quad \text{and} \quad \lim_{z \rightarrow \pm\infty} G(z; \gamma, c) \equiv 1 \quad (3)$$

where $\mathbf{1}_c$ denotes the indicator function being one only at the value c . This particular shape of G is motivated by the application of modeling real exchange rates. In addition to the given interpretation in the Introduction of getting pulled back to one regime when the process drifts off too much, we require symmetry as the exchange rate between for example US Dollars and Euros should be modeled with the same setup as the exchange rate between Euros and US Dollars.

General STAR models have not yet been studied systematically, if possible at all. As one often chooses $p = 1$ in practical applications, most of the results stated below are special cases with respect to the choice of the parameters.

2.1 The ESTAR-Model

The Exponential STAR(1) model is defined by choosing

$$G(z; \gamma, c) = 1 - \exp(-\gamma(z - c)^2), \quad z \in \mathbb{R}, \quad (4)$$

which results according to (2) for $d = p = 1$, $\psi_0 = \varphi_0 = 0$, $c = 0$, $\psi = \psi_1$ and $\varphi = \varphi_1$ in

$$y_t = [\psi + \varphi \exp(-\gamma y_{t-1}^2)] y_{t-1} + \varepsilon_t, \quad t \geq 1, \quad (5)$$

with initial variable y_0 .

Example 2.2.

Figures 1 and 2 show two realizations of the ESTAR process (5) of length $T = 500$, both generated with $\psi = 1$ and $\varphi = -0.45$. The variances are chosen as $\sigma^2 = 4$ and $\sigma^2 = 0.2^2$, respectively.

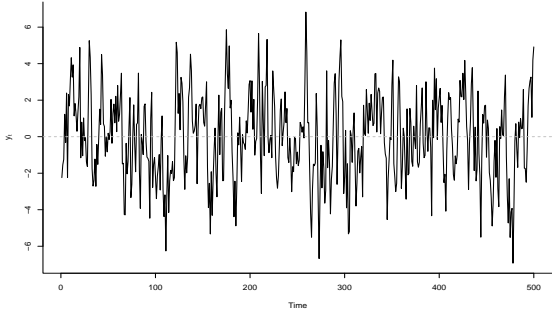


Figure 1: $\sigma = 2, \psi = 1$ and $\varphi = -0.45$.

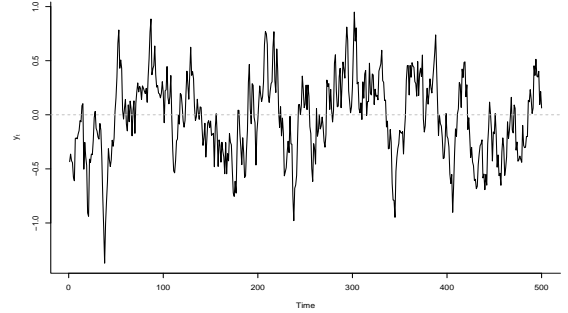


Figure 2: $\sigma = 0.2, \psi = 1$ and $\varphi = -0.45$.

As only very little is known about the theoretical properties of an ESTAR process we state a result about the moments, which will also play an important role in the following section and is proven in the Appendix.

Lemma 2.3 (Moments of y_t).

Let y_t be as in (5) where y_0 and $\varepsilon_t, t \geq 1$, have a density that is symmetric around 0. Then, for $n, k \in \mathbb{N}_0$ and all $t \geq 0$,

$$\mathbb{E} \left[\exp(-n\gamma y_t^2) y_t^{2k+1} \right] = 0, \quad (6)$$

which means in particular that all odd moments of y_t vanish for all $t \geq 0$ by choosing $n = 0$. Moreover, if we assume that y_t is normally distributed together with $|\psi| + |\varphi| < 1$ we have

$$\lim_{\sigma^2 \rightarrow 0} \mathbb{E} \left[\exp(-n\gamma y_t^2) y_t^{2k} \right] = \sigma^{2k} (1 + o_P(\sigma^2)) \quad (7)$$

for all $t \geq 0$ and $k \in \mathbb{N}_0$. The above statement yields a rate for the even moments tending to zero by setting again $n = 0$.

2.2 The Identification Problem

If transition functions G_1 and G_2 , resulting from different parameter combinations, cannot be distinguished, it is obviously nearly impossible to fit a ‘good’ model to given data. Whenever changes of the parameters do not result in significant changes of the transition function, we speak of the so-called identification problem. Due to (5) this happens in the ESTAR setting for extreme values (i.e. large values or values close to zero) of γy_{t-1}^2 , caused either by γ or by y_{t-1}^2 . The latter turns out to occur for very small or very large values of the error term variance σ^2 . This observation

is clearly in contrast to linear models and was mentioned by [Luukkonen et al. \(1988\)](#). However, to the best of our knowledge it has not been further considered systematically in the literature since.

Before giving more profound results, we want to describe the intuition behind the identification problem. With respect to γ , it is obvious that for different large values (say roughly $\gamma > 1$), the corresponding transition functions hardly change any more. As for y_t^2 , one can already see from [Figures 1 and 2](#), which show realizations of ESTAR processes with different values of σ^2 , that -while only appearing implicitly in the definition (5)- the error term variance influences the behavior of the process. Large values for σ^2 allow the error term to dominate the process, resulting in large values for y_t and causing the identification problem, independent of the choice of γ . On the other hand, very small values for σ^2 result in small values of y_t .

Consider parameters γ and σ^2 for which the identification problem is present. As a consequence, the transition function G in the ESTAR model is either close to zero or close to one. This means that one of the two regimes is no longer present. The transition parameter γ as well as one of the autoregressive parameters are therefore unidentified and can not be estimated consistently. To illustrate this behavior, we estimate the parameter vector $(\psi, \varphi, \gamma, \sigma^2)$ by means of the conditional Maximum Likelihood method. The resulting highly biased estimators for γ using different choices of σ are summarized in [Table 1](#).

| $\gamma \backslash \sigma$ | | 0.1 | | 0.5 | | 1.0 | |
|----------------------------|-----------------|--------|-------|--------|-------|--------|---------|
| | | Mean | SD | Mean | SD | Mean | SD |
| 0.2 | $\hat{\psi}$ | 0.755 | 0.008 | 0.781 | 0.084 | 0.774 | 0.099 |
| | $\hat{\varphi}$ | -0.432 | 0.040 | -0.605 | 0.337 | -0.581 | 0.224 |
| | $\hat{\gamma}$ | 0.334 | 0.149 | 0.698 | 1.425 | 0.274 | 0.303 |
| | $\hat{\sigma}$ | 0.100 | 0.000 | 0.499 | 0.013 | 0.997 | 0.025 |
| 0.7 | $\hat{\psi}$ | 0.755 | 0.008 | 0.778 | 0.096 | 0.773 | 0.150 |
| | $\hat{\varphi}$ | -0.428 | 0.044 | -0.604 | 0.237 | -0.520 | 0.157 |
| | $\hat{\gamma}$ | 0.828 | 0.148 | 0.988 | 1.092 | 0.867 | 0.852 |
| | $\hat{\sigma}$ | 0.100 | 0.000 | 0.499 | 0.013 | 0.998 | 0.024 |
| 1.0 | $\hat{\psi}$ | 0.755 | 0.008 | 0.774 | 0.106 | 0.783 | 0.168 |
| | $\hat{\varphi}$ | -0.429 | 0.043 | -0.581 | 0.208 | -0.516 | 0.166 |
| | $\hat{\gamma}$ | 1.136 | 0.149 | 1.279 | 1.245 | 5.642 | 186.820 |
| | $\hat{\sigma}$ | 0.100 | 0.000 | 0.499 | 0.012 | 0.998 | 0.025 |
| 1.3 | $\hat{\psi}$ | 0.754 | 0.007 | 0.775 | 0.113 | 0.783 | 0.185 |
| | $\hat{\varphi}$ | -0.427 | 0.044 | -0.560 | 0.185 | -0.510 | 0.180 |
| | $\hat{\gamma}$ | 1.436 | 0.149 | 1.586 | 1.346 | 1.849 | 5.433 |
| | $\hat{\sigma}$ | 0.100 | 0.000 | 0.499 | 0.012 | 0.998 | 0.026 |

Table 1: Estimation results for ESTAR: $y_t = 0.75y_{t-1} - 0.45y_{t-1}G(\cdot) + \varepsilon_t$

It is definitely worth studying this phenomena as it is in particular counter intuitive that tiny error term variances do not allow for good estimators as one would expect to observe (and estimate) the process well. Moreover, although not called identification problem, people are aware of the problems and a lot of subjective ‘tricks’ have been proposed and used to circumvent them, allowing for a broader range for γ and σ without experiencing unidentified parameters. The common idea is to exclude γ from the estimation process and use an alternative way to fit the model. [Haggan](#)

and Ozaki (1981) propose, for instance, to define a grid for γ and estimate only the remaining parameters, followed by a search for the best γ . By doing so, they do not estimate the transition variable γ . In order to reduce the influence of σ , Lütkepohl and Krätzig (2004, p.229) standardize the exponent present in G by writing

$$G(y_t; \gamma, c) = 1 - \exp(-\gamma y_t^2) = 1 - \exp\left(-\gamma \hat{\sigma}^2 \cdot \left(\frac{y_t^2}{\hat{\sigma}^2}\right)\right),$$

where $\hat{\sigma}$ is the standard deviation, in order to obtain a scale free γ . However, this is not the case as the resulting Volterra series (see Priestley, 1988, p.25) is not bounded.

Although these modifications seem to help in certain situations they are not quite satisfying as it is hard to reproduce the parameter estimates and as they have not been studied well mathematically. However, it would indeed be desirable to have a mathematical unified approach for the estimation problem, in particular for very small σ^2 , as one does find tiny estimated values $\hat{\sigma}$ in practical applications. See for example Gatti et al. (1998, p.56) or Öcal (2000, p.129), where small values for σ^2 together with huge estimates for γ are computed.

We close this section by proving that for small σ^2 one indeed will never find a good estimator for the unidentified γ . Tjøstheim (1986) derives in Theorem 3.2 asymptotic normality for a conditional Maximum Likelihood estimator $\hat{\beta}$ of $\beta = (\psi, \varphi, \gamma)$ of a more general models than studied in this text. Specifying that result for the ESTAR model stated in (5) we obtain the following theorem, proved in the Appendix.

Theorem 2.4 (Asymptotic Variance of $\hat{\beta}$).

Let y_t be as in (5) where $\gamma > 0$ and where ψ and φ are chosen such that $|\psi| + |\varphi| < 1$. Let $\beta = (\psi, \varphi, \gamma)$ be the parameter vector estimated by the conditional ML estimator $\hat{\beta} = (\hat{\psi}, \hat{\varphi}, \hat{\gamma})$. Assume that $y_0 \sim N(0, \sigma^2)$ and that $\varepsilon_t \sim N(0, \sigma^2)$ for all $t \geq 1$ in addition to satisfying Assumption 2.1. Then

$$\lim_{\sigma \downarrow 0} \text{Var}(\hat{\gamma}) \rightarrow \infty. \tag{8}$$

Remarks.

- We are aware that the limiting situation in (8) never occurs in practical applications. However, the result should be read that the transition parameter γ can hardly be identified for very small sizes of the error variance, which results in biased estimators if no other correction is included in the optimization routine for deducing $\hat{\beta}$.
- The condition $|\psi| + |\varphi| < 1$ is only included in order to apply Theorem 3.2 of Tjøstheim (1986).
- We restrict the parameter vector in Theorem 2.4 to the three dimensional β not containing σ^2 only for technical reasons. In Tjøstheim (1986, Theorem 5.2) one can also find a general limiting result for $\tilde{\beta} = (\psi, \varphi, \gamma, \sigma)$. That however neither yields any new information about the behavior of $\hat{\gamma}$, nor any substantial information about the remaining parameters and has therefore not been included in order to keep the proof of the above theorem somewhat readable.
- Theorem 2.4 only covers the case $\sigma \rightarrow 0$. As mentioned earlier, $\sigma \rightarrow \infty$ causes the identification problem, too. This is not just intuitive but has also been supported by simulation studies. Details are not included here as small values for σ^2 are the more interesting case in practical applications.

3 The T-STAR Model

In the ESTAR model, unidentified parameters occur for values (γ, σ) in a certain region, say $R_{\gamma, \sigma}^E$. In particular estimating γ in the presence of a small $\sigma \in R_{\gamma, \sigma}^E$ becomes impossible while on the other hand those values for σ^2 are used in the literature.

We now propose to choose a different, new transition function within the STAR-framework, resulting in the so-called T-STAR model. The identification problem is also present in this new model. However, $R_{\gamma, \sigma}^T$, the region for which the identification of the parameters is not possible, seems to be smaller than $R_{\gamma, \sigma}^E$. Hence, a direct estimation of the parameters for more extreme values of (γ, σ) is possible, making the T-STAR model superior to the ESTAR-model.

In Section 3.1 the T-STAR model is defined. A linearity and a unit root test are derived in Sections 3.2 and 3.3, respectively. Section 3.4 then gives an overview of the performed Monte Carlo Simulations.

3.1 The Model and Estimators

Motivated by the relation of the transition function (4) to the normal density function, we now define the T-STAR model, by proposing a transition function similar to the density of Student's t-distribution, i.e.

$$G(z; \gamma, c) = \left[1 - (1 + (z - c)^2)^{-\gamma} \right], \quad z \in \mathbb{R}, \quad (9)$$

with $\gamma > 0$, $1 \leq d \leq p$ and $c \in \mathbb{R}$. The parameters γ and c can be interpreted just as in the ESTAR model as transition and location variable. Also, properties like boundedness, the limit behavior for $z \rightarrow \pm\infty$ and $\gamma \rightarrow \pm\infty$ (see (3)) as well as the shape of G remain unchanged compared to (4). The T-STAR model can therefore be seen as an alternative model to the ESTAR model, applicable to the same situations.

The present identification problem causes less problems as different functions G are clearly distinct for a larger range of values for γ than in the ESTAR model. This is also visible in Figure 3 which illustrates for different values of γ the resulting transition functions in comparison to the ESTAR setting shown in Figure 4 (note the different scale on the x-axis).

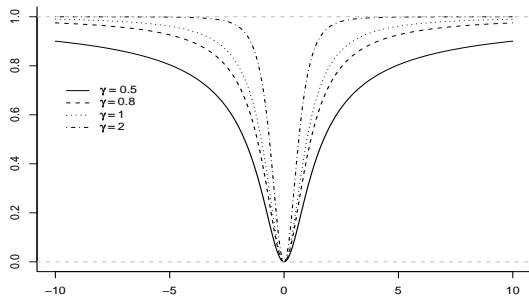


Figure 3: T-STAR: Transition function for different γ .

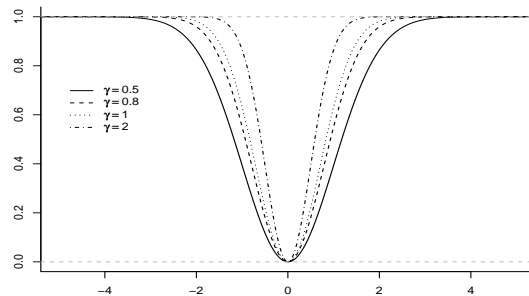


Figure 4: ESTAR: Transition function for different γ .

As γ and y_t^2 no longer appear as a product in the transition function (9), the interplay between these two parameters is reduced. For different values of γ the transition functions still differ even

for small values of σ^2 . This however implies that better estimates for $\beta = (\psi, \varphi, \gamma, \sigma^2)$ are obtained even for parameter combinations of γ and σ^2 that cause problems in the ESTAR setting. For this compare Table 1 and Table 2, where the latter displays the performance of the estimator $\hat{\gamma}$ in the T-STAR model.

| $\sigma \backslash \gamma$ | | 0.3 | | 0.5 | | 1.0 | |
|----------------------------|-----------------|-------|-------|-------|-------|-------|-------|
| | | Mean | SD | Mean | SD | Mean | SD |
| 0.2 | $\hat{\psi}$ | 0.300 | 0.001 | 0.300 | 0.002 | 0.152 | 0.374 |
| | $\hat{\varphi}$ | 0.400 | 0.001 | 0.400 | 0.001 | 0.746 | 0.119 |
| | $\hat{\gamma}$ | 0.200 | 0.007 | 0.200 | 0.016 | 0.640 | 1.031 |
| | $\hat{\sigma}$ | 0.300 | 0.009 | 0.500 | 0.006 | 0.998 | 0.023 |
| 0.8 | $\hat{\psi}$ | 0.300 | 0.004 | 0.300 | 0.005 | 0.107 | 0.355 |
| | $\hat{\varphi}$ | 0.400 | 0.001 | 0.400 | 0.001 | 0.744 | 0.136 |
| | $\hat{\gamma}$ | 0.800 | 0.006 | 0.800 | 0.007 | 1.125 | 1.214 |
| | $\hat{\sigma}$ | 0.300 | 0.009 | 0.500 | 0.006 | 0.998 | 0.022 |
| 1.0 | $\hat{\psi}$ | 0.300 | 0.004 | 0.300 | 0.005 | 0.119 | 0.340 |
| | $\hat{\varphi}$ | 0.400 | 0.001 | 0.400 | 0.001 | 0.620 | 0.152 |
| | $\hat{\gamma}$ | 1.000 | 0.007 | 1.000 | 0.006 | 1.303 | 1.359 |
| | $\hat{\sigma}$ | 0.300 | 0.10 | 0.500 | 0.006 | 0.998 | 0.022 |
| 1.3 | $\hat{\psi}$ | 0.300 | 0.004 | 0.300 | 0.007 | 0.148 | 0.302 |
| | $\hat{\varphi}$ | 0.400 | 0.001 | 0.400 | 0.001 | 0.581 | 0.168 |
| | $\hat{\gamma}$ | 1.300 | 0.004 | 1.300 | 0.005 | 1.583 | 1.653 |
| | $\hat{\sigma}$ | 0.300 | 0.010 | 0.500 | 0.006 | 0.998 | 0.022 |

Table 2: Estimation results for T-STAR: $y_t = 0.3y_{t-1} + 0.4y_{t-1}\mathcal{G}(\cdot) + \varepsilon_t$

3.2 Linearity testing

The procedure we derive in this section for testing linearity against non-linear T-STAR dynamics is related to the test against non-linearity proposed by Luukkonen et al. (1988). The transition function G is first approximated by a suitable linear function; a common practice in non-linear time series analysis (see also Teräsvirta, 1984). Afterwards, a simple F -test is performed.

For constructing the test it is convenient to use representation (2), i.e.

$$y_t = [\Psi w_t] + [\Phi w_t] \times G(y_{t-d}; \gamma, c) + \varepsilon_t. \quad (10)$$

Under linearity the autoregressive parameters are then identical for both regimes. Thus, we do not have to allow for switching between identical regimes and achieve a more parsimonious model by using a linear AR(p) model.

The pair of hypothesis we are interested in can be expressed either as

$$H_0 : \Phi = \mathbf{0}_{(1 \times p)} \quad \text{vs.} \quad H_1 : \text{at least one } \varphi_i \neq 0; \quad i = 1, \dots, p$$

or

$$H_0 : \gamma = 0 \quad \text{vs.} \quad H_1 : \gamma > 0.$$

In both cases the T-STAR model (10) reduces to a linear autoregressive model of order p . However, our test procedure employs the former pair of hypothesis.

Under H_0 the alternative is not identified, given that the vector Φ and c can take on any value without changing the value of the likelihood function when $\gamma = 0$ and vice versa. This can be circumvented by replacing G with a linear approximation. Based on the Binomial series, i.e.

$$(1+x)^{-m} = 1 + \sum_{n=1}^{\infty} (-1)^n \frac{m(m+1)(m+2)\dots(m+n-1)}{n!} x^n, \quad m > 0, \quad (11)$$

the transition function G in (10) can be approximated arbitrarily well by

$$G_k(\cdot) = \sum_{n=1}^k (-1)^n \frac{\gamma(\gamma+1)\dots(\gamma+n-1)(y_{t-d}-c)^{2n}}{n!} \quad (12)$$

choosing $x = (y_{t-d}-c)^2$ and $m = \gamma$ in (11) as well as a suitable k . After expanding the terms $(y_{t-d}-c)^{2n}$, $n = 1, \dots, k$, and some rearrangements, we obtain the auxiliary regression model for a fixed $d \leq p$ and k

$$y_t = \sum_{i=1}^p \phi_i y_{t-i} + \sum_{j=1}^p \delta_j^0 y_{t-j} + \sum_{j=1}^p \delta_j^1 y_{t-j} y_{t-d} + \sum_{j=1}^p \delta_j^2 y_{t-j} y_{t-d}^2 + \dots + \sum_{j=1}^p \delta_j^{2k} y_{t-j} y_{t-d}^{2k} + u_t. \quad (13)$$

If the location parameter is a priori restricted to $c = 0$ then the model simplifies as only the odd powers of y_{t-} remain. The error terms in the regression are now denoted by u_t rather than ε_t as they are the sum of the original error terms and the approximation error caused by replacing G with G_k .

A test against non-linearity can then be carried out using a simple F -test for a subvector of parameters. Under the null the actual model is linear and hence the approximation error is zero leading to $u_t = \varepsilon_t$. Consequently the properties of the error term under the null and thus the asymptotic distribution of the F -test remain unaffected.

Example 3.1.

As an example consider the simple T-STAR(1) model from above,

$$y_t = \psi_1 y_{t-1} + \varphi_1 y_{t-1} \left[1 - (1 + (y_{t-d} - c)^2)^{-\gamma} \right] + \varepsilon_t,$$

with nonzero location parameter c . Approximating G by G_3 first results in

$$y_t = \psi_1 y_{t-1} + \varphi_1 y_{t-1} \left[\gamma(y_{t-1} - c)^2 - \frac{1}{2}\gamma(\gamma+1)(y_{t-1} - c)^4 + \frac{1}{6}\gamma(\gamma+1)(\gamma+2)(y_{t-1} - c)^6 \right] + u_t.$$

Using

$$\begin{aligned} (y_{t-1} - c)^2 &= y_{t-1}^2 - 2y_{t-1}c + c^2, \\ (y_{t-1} - c)^4 &= y_{t-1}^4 - 4y_{t-1}^3c + 6y_{t-1}^2c^2 - 4y_{t-1}c^3 + c^4, \\ (y_{t-1} - c)^6 &= y_{t-1}^6 - 6y_{t-1}^5c + 15y_{t-1}^4c^2 - 20y_{t-1}^3c^3 + 15y_{t-1}^2c^4 - 6y_{t-1}c^5 + c^6, \end{aligned}$$

we then obtain the auxiliary regression model (see (13))

$$y_t = \phi_1 y_{t-1} + \delta_1^0 y_{t-1} + \delta_1^1 y_{t-1}^2 + \delta_1^2 y_{t-1}^3 + \delta_1^3 y_{t-1}^4 + \delta_1^4 y_{t-1}^5 + \delta_1^5 y_{t-1}^6 + \delta_1^6 y_{t-1}^7 + u_t$$

where

$$\begin{aligned}
\delta_1^0 &= \varphi_1 \gamma c^2 + \frac{1}{6} \varphi_1 \gamma (\gamma + 1) (\gamma + 2) c^6, \\
\delta_1^1 &= -2c \varphi_1 \gamma + 2 \varphi_1 \gamma (\gamma + 1) - \varphi_1 \gamma (\gamma + 1) (\gamma + 2) c^5, \\
\delta_1^2 &= \varphi_1 \gamma - 3 \varphi_1 \gamma (\gamma + 1) c^2 + \frac{5}{2} \varphi_1 \gamma (\gamma + 1) (\gamma + 2) c^4, \\
\delta_1^3 &= 2 \varphi_1 \gamma (\gamma + 1) c - \frac{1}{2} \varphi_1 \gamma (\gamma + 1) c^4 - \frac{10}{3} \varphi_1 \gamma (\gamma + 1) (\gamma + 2) c^3, \\
\delta_1^4 &= -\frac{1}{2} \varphi_1 \gamma (\gamma + 1) + \frac{5}{2} \varphi_1 \gamma (\gamma + 1) (\gamma + 2) c^2, \\
\delta_1^5 &= -\varphi_1 \gamma (\gamma + 1) (\gamma + 2) c, \\
\delta_1^6 &= \frac{1}{6} \varphi_1 \gamma (\gamma + 1) (\gamma + 2).
\end{aligned}$$

The hypothesis of linearity against T-STAR can be tested via an F-test for the null

$$H_0 : \delta_1^0 = \dots = \delta_1^6 = 0 \quad \text{vs.} \quad H_1 : \text{at least one } \delta_1^i \neq 0; \quad i = 1, \dots, 6.$$

Monte Carlo Simulations for this example can be found in Section 3.4.

3.3 Unit Root Testing

Kapetanios et al. (2003) develop a unit root test in the ESTAR framework and compute a Dickey-Fuller type t -test in this set-up based on a first order Taylor expansion. Our test is of the same type and thus we test the null of a linear unit root process against a globally stationary T-STAR process containing a partial unit root in one regime.

Let the T-STAR(1) process again be parametrized by

$$y_t = \psi_1 y_{t-1} + \varphi_1 y_{t-1} \left[1 - (1 + y_{t-1}^2)^{-\gamma} \right] + \varepsilon_t. \quad (14)$$

Setting the location parameter c equal to zero is motivated by simulation results in Kruse (2009) that show convincing power results even if the location parameter c is set to zero ex-ante. This is also consistent with Kapetanios et al. (2003). Hence, for the sake of simplicity we constrain ourselves to this case and further impose $d = 1$ which is in line with empirical applications of non-linear time series models (see e.g. Taylor et al. (2001) or Rapach and Wohar (2006)).

The model in (14) can be written in first differences as

$$\Delta y_t = \beta y_{t-1} + \varphi_1 y_{t-1} \left[1 - (1 + y_{t-1}^2)^{-\gamma} \right] + \varepsilon_t \quad (15)$$

where $\beta = \psi_1 - 1$. Setting $\beta = 0$ yields a unit root in the first regime and we have to distinguish between two cases:

- (i) $\beta = 0$ and $\gamma > 0$: In this case we have a globally stationary T-STAR process that contains a partial unit root in the first regime, provided that $-2 < \varphi_1 < 0$ as we will assume henceforth.
- (ii) $\beta = 0$ and $\gamma = 0$: In this case the model reduces to a linear random walk.

Thus we will test case (ii) against case (i) and formulate the pair of hypotheses as

$$H_0 : \gamma = 0 \quad \text{vs.} \quad H_1 : \gamma > 0 . \quad (16)$$

We now proceed in the same way as in the previous section and approximate the nonlinearity with a Binomial expansion as in (12) setting the number of summands to $k = 3$. This yields the auxiliary regression (see also (13)):

$$\Delta y_t = \delta_1^2 y_{t-1}^3 + \delta_1^4 y_{t-1}^5 + \delta_1^6 y_{t-1}^7 + u_t \quad (17)$$

so that we want to test

$$H_0 : \delta_1^2 = \delta_1^4 = \delta_1^6 = 0 \quad \text{vs.} \quad H_1 : \text{at least one } \delta_1^i \neq 0; \quad i = 2, 4, 6 .$$

As we are now dealing with three parameter restrictions we cannot use the conventional t -statistic. Therefore, an F -statistic for the significance of the whole parameter vector $\beta = (\delta_1^2, \delta_1^4, \delta_1^6)'$ needs to be computed. For that, it is convenient to write the null as

$$H_0 : R\beta = r$$

with

$$R = I_3 = \begin{bmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{bmatrix} \quad \text{and} \quad r = (0, 0, 0)' .$$

Using $\hat{r} = R\hat{\beta}$ we can write the F -statistic as

$$\begin{aligned} F^* &= \frac{1}{3}(\hat{r} - r)' [\hat{\sigma}^2 R(X'X)^{-1}R']^{-1} (\hat{r} - r) = \frac{1}{3}(R\hat{\beta})' [\hat{\sigma}^2 R(X'X)^{-1}R']^{-1} (R\hat{\beta}) \\ &= \frac{1}{3}\hat{\beta}' [\hat{\sigma}^2 (X'X)^{-1}]^{-1} \hat{\beta} \end{aligned} \quad (18)$$

where X is a $(T \times 3)$ design matrix with its t -th row given by $x_t = (y_{t-1}^3, y_{t-1}^5, y_{t-1}^7)$ and $\hat{\sigma}^2 = \frac{1}{T-4} \sum_{t=1}^T \left(\Delta y_t - \hat{\delta}_1^2 y_{t-1}^3 - \hat{\delta}_1^4 y_{t-1}^5 - \hat{\delta}_1^6 y_{t-1}^7 \right)^2$. The limit distribution of F^* under H_0 is computed in the next theorem. In the sequel we denote weak convergence by \Rightarrow and convergence in probability by \xrightarrow{P} . And a standard Brownian Motion is denoted by $B(r)$.

Theorem 3.2.

Consider the T -STAR model (14) and let ε_t satisfy Assumption 2.1. Then the test statistic F^* as given in (18) behaves asymptotically for $T \rightarrow \infty$ under the null of a random walk as follows

$$F^* \Rightarrow \frac{1}{3\sigma^2} \mathbf{v}' \mathbf{Q}^{-1} \mathbf{v},$$

where the matrices \mathbf{Q} and \mathbf{v} are given by

$$\mathbf{Q} = \begin{bmatrix} \sigma^6 \int_0^1 B^6(r) dr & \sigma^8 \int_0^1 B^8(r) dr & \sigma^{10} \int_0^1 B^{10}(r) dr \\ 0 & 0 & 0 \\ \sigma^8 \int_0^1 B^8(r) dr & \sigma^{10} \int_0^1 B^{10}(r) dr & \sigma^{12} \int_0^1 B^{12}(r) dr \\ 0 & 0 & 0 \\ \sigma^{10} \int_0^1 B^{10}(r) dr & \sigma^{12} \int_0^1 B^{12}(r) dr & \sigma^{14} \int_0^1 B^{14}(r) dr \\ 0 & 0 & 0 \end{bmatrix}$$

and

$$\mathbf{v} = \left[\sigma^4 \left\{ \frac{1}{4} \int_0^1 B(1)^4 - \frac{3}{2} \int_0^1 B(r) dr \right\} \quad \sigma^6 \left\{ \frac{1}{6} \int_0^1 B(1)^6 - \frac{5}{2} \int_0^1 B(r) dr \right\} \quad \sigma^8 \left\{ \frac{1}{8} \int_0^1 B(1)^8 - \frac{7}{2} \int_0^1 B(r) dr \right\} \right].$$

Under the alternative the test is consistent.

In order to deal with deterministic components such as non-zero intercept terms or linear trends one can use a two-step approach and de-mean or de-trend the data prior to computing the test statistic F^* . In this case the true data generating process is given by $y_t = \omega' z_t + x_t$ where $x_t = y_{t-1} + \varepsilon_t$ and ω' is a parameter vector of suitable dimensions and $z_t = 1$ for all t for the de-meanded case and $z_t = [1, t]$ for the de-trended case. The test can then be based on the OLS residuals \hat{x}_t , where the asymptotic distribution now depends on functionals of de-meanded and de-trended Brownian motion, respectively. These are given by

$$B(r) - \int_0^1 B(r) dr$$

for the de-meanded Brownian motion and by

$$B(r) + (6r - 4) \int_0^1 B(r) dr + (12r - 6) \int_0^1 r B(r) dr$$

for the de-trended Brownian motion.

Considering the case of serially correlated errors and assuming that the dependence enters in a linear fashion we can generalize our results by augmenting the auxiliary regression with lagged differences as in [Dickey and Fuller \(1979\)](#) and [Said and Dickey \(1984\)](#). The test regression then reads

$$\Delta y_t = \delta_1^2 y_{t-1}^3 + \delta_1^4 y_{t-1}^5 + \delta_1^6 y_{t-1}^7 + \sum_{i=1}^p \rho_i \Delta y_{t-i} + u_t. \quad (19)$$

The pair of hypotheses as well as the test statistic in this more general set up do not change with respect to the auxiliary regression in (17).

Theorem 3.3.

Consider the test statistic F^* as in Theorem 3.2 but computed from (19). Under the null of a unit root the test statistic maintains the same asymptotic distribution as in Theorem 3.2. Under the alternative the test statistic is consistent.

Theorem 3.3 holds also true for the case of including deterministic terms as in auxiliary regression (19). The asymptotic distribution in this case is such as in Theorem 3.2 when deterministic terms are included, i.e. replacing the standard Brownian motion with the de-meanded or de-trended Brownian motion, respectively.

Setting the approximation of the infinite sum from the Binomial series expansion to $k = 1$ it is readily seen that $\sqrt{F^*}$ has the same asymptotic distribution as the unit root test against ESTAR developed by [Kapetanios et al. \(2003\)](#) and thus the statistic F^* contains their test as a special case. It is also noteworthy that the F -test version of the ESTAR unit root test of [Kapetanios et al.](#)

(2003) that would result if the location parameter c is not set equal to zero a priori is also a special case of our test. Setting the series expansion again to $k = 1$ and also letting the location parameter $c \neq 0$ the resulting limiting distribution of the test statistic from the related auxiliary regression is the same as for the respective ESTAR unit root test and thus we also contain this test version as a special case.

Containing these tests as special cases we expect a satisfying performance also against ESTAR processes but higher power against globally stationary alternatives than the Kapetanios et al. (2003) test as indicated by a faster rate of convergence in Theorem 3.2.

3.4 Monte Carlo Simulations

In this section we study the finite sample performance of the two tests developed above for the T-STAR(1) model. Starting with the linearity test (see Section 3.2) the empirical size and power under two different data generating processes are investigated. In order to keep the experiment simple we conduct the simulations under the null of linearity using a simple AR(1) model and compute the auxiliary regression using the location parameter $c = 0$ for the first scenario and $c = 1$ for the second scenario. For all experiments reported here we perform $M = 50000$ replications combined with different sample sizes. An exception is the simulation to obtain the asymptotic critical values for the unit root test for which the sample size is set to $T = 10000$ and the number of replications to $M = 1000000$. For all power simulations reported here we consider $\alpha = 0.05$ to save space.¹

The empirical size results for the linearity test for scenario one and scenario two are stated in Tables 3 and 4, respectively. The test shows only minor deviations from its nominal level and it tends to under-reject somewhat. However, the test -although conservative- seems to be properly sized for reasonable sample sizes encountered in monthly or daily data.

| T | $\psi_1 = 0.3$ | | | $\psi_1 = 0.5$ | | | $\psi_1 = 0.8$ | | |
|------|----------------|----------------|-----------------|----------------|----------------|-----------------|----------------|----------------|-----------------|
| | $\alpha = 1\%$ | $\alpha = 5\%$ | $\alpha = 10\%$ | $\alpha = 1\%$ | $\alpha = 5\%$ | $\alpha = 10\%$ | $\alpha = 1\%$ | $\alpha = 5\%$ | $\alpha = 10\%$ |
| 100 | 0.826 | 4.258 | 8.752 | 0.738 | 4.152 | 8.440 | 0.856 | 4.194 | 8.606 |
| 200 | 0.872 | 4.442 | 9.038 | 0.786 | 4.200 | 8.664 | 0.820 | 4.206 | 8.504 |
| 500 | 0.910 | 4.546 | 9.404 | 0.876 | 4.470 | 8.986 | 0.744 | 4.248 | 8.680 |
| 1000 | 0.858 | 4.594 | 9.554 | 0.840 | 4.360 | 9.008 | 0.878 | 4.152 | 8.578 |
| 5000 | 0.886 | 4.766 | 9.906 | 0.856 | 4.536 | 9.354 | 0.888 | 4.498 | 8.986 |

Table 3: Size results for scenario one.

| T | $\psi_1 = 0.3$ | | | $\psi_1 = 0.5$ | | | $\psi_1 = 0.8$ | | |
|------|----------------|----------------|-----------------|----------------|----------------|-----------------|----------------|----------------|-----------------|
| | $\alpha = 1\%$ | $\alpha = 5\%$ | $\alpha = 10\%$ | $\alpha = 1\%$ | $\alpha = 5\%$ | $\alpha = 10\%$ | $\alpha = 1\%$ | $\alpha = 5\%$ | $\alpha = 10\%$ |
| 100 | 0.912 | 4.298 | 8.608 | 0.818 | 4.256 | 8.594 | 0.838 | 4.402 | 8.744 |
| 200 | 0.776 | 4.200 | 8.764 | 0.802 | 3.986 | 8.268 | 0.796 | 4.150 | 8.504 |
| 500 | 0.818 | 4.266 | 8.866 | 0.776 | 4.184 | 8.434 | 0.892 | 4.270 | 8.530 |
| 1000 | 0.882 | 4.626 | 9.276 | 0.868 | 4.318 | 8.708 | 0.848 | 4.218 | 8.528 |
| 5000 | 0.968 | 4.584 | 9.372 | 0.882 | 4.450 | 8.892 | 0.792 | 4.260 | 8.764 |

Table 4: Size results for scenario two.

¹The results for the cases $\alpha = 0.01$ and $\alpha = 0.1$ as well as all other unreported results are available from the authors upon request.

Turning to the power experiments we use a T-STAR(1) model and several values for ψ_1 , φ_1 and γ . The results are shown in Table 5 for scenario one and in Table 6 for scenario two. The results suggest that the linearity test is a useful device to detect non-linearity in the data. As expected the rejection frequency becomes closer to 100% the more pronounced the difference between the regimes is and/or the larger the sample size is. Overall we obtain very similar power results against T-STAR compared to Luukkonen et al. (1988) for their linearity test against ESTAR. Unreported experiments confirmed that the proposed linearity test has also similar high power against the other non-linear alternatives ESTAR, LSTAR and Double LSTAR (see Jansen and Teräsvirta, 1996). Reasonable power results were also obtained against the Markov switching model proposed by Hamilton (1989).

| | | T = 200 | | | T = 500 | | | T = 1000 | | |
|----------|-------------|----------|--------|--------|---------|--------|--------|----------|--------|--------|
| | | γ | 0.5 | 0.8 | 1.0 | 0.5 | 0.8 | 1.0 | 0.5 | 0.8 |
| ψ_1 | φ_1 | | | | | | | | | |
| 0.3 | 0.5 | 5.452 | 5.468 | 5.650 | 7.502 | 8.628 | 8.708 | 11.796 | 14.488 | 15.014 |
| 0.3 | 0.6 | 6.638 | 7.414 | 7.686 | 12.534 | 16.060 | 16.154 | 24.166 | 32.020 | 31.766 |
| 0.3 | 0.7 | 9.434 | 11.456 | 11.570 | 22.000 | 29.264 | 29.014 | 44.362 | 57.616 | 58.062 |
| 0.3 | 0.8 | 13.642 | 17.872 | 18.238 | 36.348 | 49.194 | 48.964 | 69.112 | 83.718 | 83.360 |
| 0.3 | 0.9 | 21.276 | 29.262 | 28.098 | 56.726 | 74.030 | 71.142 | 89.640 | 97.252 | 96.482 |

Table 5: Power results for scenario one.

| | | T = 200 | | | T = 500 | | | T = 1000 | | |
|----------|-------------|----------|--------|--------|---------|--------|--------|----------|--------|--------|
| | | γ | 0.5 | 0.8 | 1.0 | 0.5 | 0.8 | 1.0 | 0.5 | 0.8 |
| ψ_1 | φ_1 | | | | | | | | | |
| 0.3 | 0.5 | 5.964 | 6.970 | 7.070 | 10.302 | 13.826 | 14.790 | 19.392 | 27.412 | 30.088 |
| 0.3 | 0.6 | 9.228 | 11.804 | 12.514 | 21.240 | 31.126 | 34.288 | 45.714 | 63.610 | 68.202 |
| 0.3 | 0.7 | 14.444 | 20.502 | 22.364 | 40.652 | 58.280 | 62.524 | 77.352 | 91.816 | 94.004 |
| 0.3 | 0.8 | 24.126 | 34.482 | 36.886 | 66.438 | 84.038 | 86.562 | 95.732 | 99.448 | 99.634 |
| 0.3 | 0.9 | 38.434 | 53.024 | 53.978 | 87.566 | 97.006 | 96.840 | 99.692 | 99.998 | 99.986 |

Table 6: Power results for scenario two.

The simulation results for the unit root test are presented below. First we report the asymptotic critical values for the unit root test in Table 7. Case 1 denotes raw data, i.e. no deterministic components, Case 2 denotes the case of de-meaned data and Case 3 denotes the case of de-trended data. The results from the size experiments are summarized in Table 8. For larger sample sizes ($T > 500$) the test is correctly sized and as the sample size increases it reaches its nominal level. For smaller sample sizes some minor size distortions are visible but the overall impression is that the test maintains good size properties also for smaller sample sizes.

For the power experiment for the unit root test we exemplarily show our results for $T = 200$ and $t = 500$ and various values for γ , ψ and φ (Tables 9 and 10). The results indicate a good overall performance of the unit root test in all sample sizes considered. The ability to distinguish between a unit root process and a globally stationary T-STAR model increases if either the difference between the regimes becomes larger or even faster if the sample size increases.

| α | Case 1 | Case 2 | Case 3 |
|----------|--------|--------|--------|
| 1% | 4.730 | 5.477 | 6.595 |
| 2.5% | 3.124 | 4.722 | 5.783 |
| 5% | 3.458 | 4.137 | 5.136 |
| 7.5% | 3.124 | 3.778 | 4.739 |
| 10% | 2.884 | 3.515 | 4.450 |

Table 7: Asymptotic critical values for unit root testing.

| T | Case 1 | | | Case 2 | | | Case 3 | | |
|-------|----------------|----------------|-----------------|----------------|----------------|-----------------|----------------|----------------|-----------------|
| | $\alpha = 1\%$ | $\alpha = 5\%$ | $\alpha = 10\%$ | $\alpha = 1\%$ | $\alpha = 5\%$ | $\alpha = 10\%$ | $\alpha = 1\%$ | $\alpha = 5\%$ | $\alpha = 10\%$ |
| 100 | 0.826 | 3.898 | 7.628 | 0.786 | 3.484 | 6.916 | 0.954 | 3.830 | 7.484 |
| 200 | 0.882 | 4.106 | 8.246 | 0.748 | 3.648 | 7.436 | 0.758 | 3.594 | 7.218 |
| 500 | 0.912 | 4.504 | 9.168 | 0.864 | 4.174 | 8.602 | 0.828 | 3.998 | 8.146 |
| 1000 | 0.932 | 4.920 | 9.528 | 0.904 | 4.370 | 9.076 | 0.862 | 4.324 | 8.868 |
| 5000 | 0.952 | 4.958 | 10.060 | 0.916 | 4.734 | 9.692 | 1.048 | 4.952 | 9.766 |
| 10000 | 0.984 | 5.036 | 10.006 | 1.004 | 4.982 | 9.990 | 0.966 | 4.918 | 9.886 |
| 50000 | 0.974 | 4.968 | 10.054 | 1.000 | 4.978 | 10.062 | 0.998 | 5.012 | 9.882 |

Table 8: Size for unit root testing using asymptotic critical values [in %].

As empirical studies using smooth transition models such as ESTAR frequently find very small variances of the innovation term we examine the behavior of the newly developed unit root test against T-STAR in such a framework. Studying this behavior is critical since Kruse et al. (2008) show via Monte Carlo simulation that under small error term variances the power of unit root tests developed for non-linear models rapidly deteriorates. We report simulation results for small sample sizes of $T = 100$ (Table 11) and $T = 500$ (Table 12) and consider error term standard deviations of $\sigma_\varepsilon = 0.1$. The results show satisfying power results even for such small sample sizes. Low power results are only found for cases in which the difference between the regimes is only very small or the transition is so slow that only little observations are in the stationary regime. In these cases it is notoriously hard to distinguish between the two regimes and as a consequences the power decreases. However, the power is still high enough to deliver reliable test results and is in particular higher than found by Kruse et al. (2008) for extant test. In the case $T = 500$ no decline in power is visible and the test works under small error variances just as well as under white noise disturbances.

The newly developed test in particular shows better power properties as the test developed by Kapetanios et al. (2003) and therefore yields more reliable results in empirical applications as indicated by Kruse at al. (2009).

With regards to the linearity test we find the power only slightly reduced under small error variances compared to the white noise assumption. These results however are unreported to save space.

| $T = 200$ | | Case 1 | | | Case 2 | | | Case 3 | | | |
|-----------|-------------|----------|--------|--------|--------|--------|--------|--------|--------|--------|--------|
| ψ_1 | φ_1 | γ | 0.5 | 0.8 | 1.0 | 0.5 | 0.8 | 1.0 | 0.5 | 0.8 | 1.0 |
| 1.0 | 0.3 | | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 |
| 1.0 | 0.4 | | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 |
| 1.0 | 0.5 | | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 |
| 1.0 | 0.6 | | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 |
| 1.0 | 0.7 | | 99.95 | 99.97 | 99.97 | 99.95 | 10.00 | 100.00 | 99.97 | 10.00 | 100.00 |
| 1.0 | 0.8 | | 95.52 | 98.86 | 99.36 | 96.58 | 99.30 | 99.59 | 97.65 | 99.46 | 99.71 |
| 1.0 | 0.9 | | 45.95 | 58.63 | 62.44 | 55.04 | 67.01 | 70.45 | 65.79 | 75.30 | 77.97 |
| 1.0 | 0.95 | | 14.90 | 18.44 | 19.47 | 23.68 | 27.84 | 28.76 | 36.00 | 39.77 | 41.77 |

Table 9: Power results for unit root testing using asymptotic critical values [in %]

| $T = 500$ | | Case 1 | | | Case 2 | | | Case 3 | | | |
|-----------|-------------|----------|--------|--------|--------|--------|--------|--------|--------|--------|--------|
| ψ_1 | φ_1 | γ | 0.5 | 0.8 | 1.0 | 0.5 | 0.8 | 1.0 | 0.5 | 0.8 | 1.0 |
| 1.0 | 0.3 | | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 |
| 1.0 | 0.4 | | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 |
| 1.0 | 0.5 | | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 |
| 1.0 | 0.6 | | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 |
| 1.0 | 0.7 | | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 |
| 1.0 | 0.8 | | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 |
| 1.0 | 0.9 | | 99.84 | 99.98 | 99.99 | 99.88 | 99.974 | 99.99 | 99.93 | 99.99 | 99.99 |
| 1.0 | 0.95 | | 74.48 | 83.15 | 85.12 | 80.29 | 87.380 | 88.96 | 85.95 | 91.19 | 92.59 |

Table 10: Power results for unit root testing using asymptotic critical values [in %]

4 Empirical Illustration

To illustrate the application of the newly introduced T-STAR model with empirical data we investigate one of the most highly debated theories in international finance: the purchasing power parity (PPP). The initial finding of a unit root in real exchange rates by [Meese and Rogoff \(1988\)](#) subsequently shifted the interest in modeling real exchange rates to non-linear models (see e.g. [Taylor et al., 2001](#)).

Technically spoken the real exchange rate should be non-linear but globally stationary and not behave like a unit root process to support PPP.

To ensure comparability we use the same data that has been analyzed by [Taylor et al. \(2001\)](#) and by [Rapach and Wohar \(2006\)](#). Namely, we analyze monthly real exchange data for Germany against the US from 1973:02 - 1996:12.² The series is depicted in Figure 5.

²The data set is available from David Rapach's website at: <http://pages.slu.edu/faculty/rapachde/Nlfit.zip>.

| $T = 100$ | | Case 1 | | | Case 2 | | | Case 3 | | | |
|-----------|-------------|----------|--------|--------|--------|--------|--------|--------|--------|--------|--------|
| ψ_1 | φ_1 | γ | 0.5 | 0.8 | 1.0 | 0.5 | 0.8 | 1.0 | 0.5 | 0.8 | 1.0 |
| 1.0 | 0.3 | | 77.108 | 95.104 | 98.174 | 79.610 | 95.388 | 98.378 | 84.538 | 96.526 | 98.716 |
| 1.0 | 0.4 | | 66.452 | 88.956 | 95.106 | 69.914 | 90.042 | 95.444 | 77.190 | 92.464 | 96.662 |
| 1.0 | 0.5 | | 52.768 | 77.878 | 87.560 | 58.408 | 80.730 | 88.988 | 67.572 | 85.324 | 91.724 |
| 1.0 | 0.6 | | 38.304 | 60.896 | 72.470 | 44.878 | 66.622 | 76.096 | 56.658 | 73.874 | 81.682 |
| 1.0 | 0.7 | | 24.404 | 40.558 | 49.668 | 33.204 | 48.180 | 57.018 | 45.100 | 59.096 | 65.972 |
| 1.0 | 0.8 | | 13.952 | 21.662 | 26.456 | 22.466 | 30.732 | 35.064 | 34.442 | 42.484 | 47.878 |
| 1.0 | 0.9 | | 6.714 | 8.604 | 9.888 | 13.770 | 16.748 | 17.924 | 24.864 | 28.176 | 29.830 |
| 1.0 | 0.95 | | 4.608 | 5.062 | 5.298 | 10.360 | 11.518 | 12.048 | 21.382 | 21.926 | 22.620 |

Table 11: Power results for unit root testing [in %] with $\sigma_\varepsilon = 0.1$

| $T = 500$ | | Case 1 | | | Case 2 | | | Case 3 | | | |
|-----------|-------------|----------|--------|--------|--------|--------|--------|--------|---------|--------|--------|
| ψ_1 | φ_1 | γ | 0.5 | 0.8 | 1.0 | 0.5 | 0.8 | 1.0 | 0.5 | 0.8 | 1.0 |
| 1.0 | 0.3 | | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.000 | 100.00 | 100.00 |
| 1.0 | 0.4 | | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.000 | 100.00 | 100.00 |
| 1.0 | 0.5 | | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.000 | 100.00 | 100.0 |
| 1.0 | 0.6 | | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.000 | 100.00 | 100.00 |
| 1.0 | 0.7 | | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.000 | 100.00 | 100.00 |
| 1.0 | 0.8 | | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 100.00 | 99.99 | 100.00 | 100.00 |
| 1.0 | 0.9 | | 93.25 | 99.09 | 99.70 | 94.01 | 99.19 | 99.74 | 95.57 | 99.38 | 99.77 |
| 1.0 | 0.95 | | 44.70 | 63.54 | 70.72 | 53.48 | 70.29 | 76.52 | 64.24 | 77.55 | 82.39 |

Table 12: Power results for unit root testing [in %] with $\sigma = 0.1$



Figure 5: Monthly log real exchange rate for Germany

We choose the lag length to be used subsequently with the Bayesian information criterion (BIC) which yields a lag length of $p = 1$. Applying the linearity test against T-STAR described in Section 3.2 we obtain a test statistic of 3.69 which is significant on the $\alpha = 5\%$ level of significance and thus we reject the null of linearity.

Validity of the PPP suggests that the real exchange rate should be a globally stationary process albeit non-linear. Applying the ESTAR unit root test developed by [Kapetanios et al. \(2003\)](#) as well as the unit root test against T-STAR developed in Section 3.3 yields support for the PPP. Both test are able to reject the null of a random walk on the $\alpha = 5\%$ level of significance. These test results support the theory that transaction costs in financial markets lead to a non-linear convergence to a long-run equilibrium and thus support the validity of the PPP as a long run concept.

Since the data has already been under study by [Taylor et al. \(2001\)](#) we adopt the parameter estimates they found and which have also been confirmed by estimations undertaken by [Rapach and Wohar \(2006\)](#).³ Table 13 shows the estimation results for the parameter γ for the model under the null for the ESTAR and the T-STAR model respectively.

| | ESTAR | T-STAR |
|----------------------------|-------|---------|
| $\hat{\gamma}$ | 0.264 | 275.284 |
| $\hat{\sigma}_\varepsilon$ | 0.035 | 0.032 |

Table 13: Estimation of γ under the null.

At a first glance the estimation result for ESTAR looks reasonable. But if we plot the estimated transition function against the transition variable y_{t-1} and against time (see Figure 6) we get to the conclusion that the ESTAR model basically reduces to a random walk model as the transition function is always close or equal to zero effectively switching off the stationary regime.

³It should be noted that [Taylor et al. \(2001\)](#) and [Rapach and Wohar \(2006\)](#) also estimated the location parameter c . However, as their estimate is very close to zero, namely $c = -0.007$, we restrict $c = 0$ in our estimation to keep it simple.

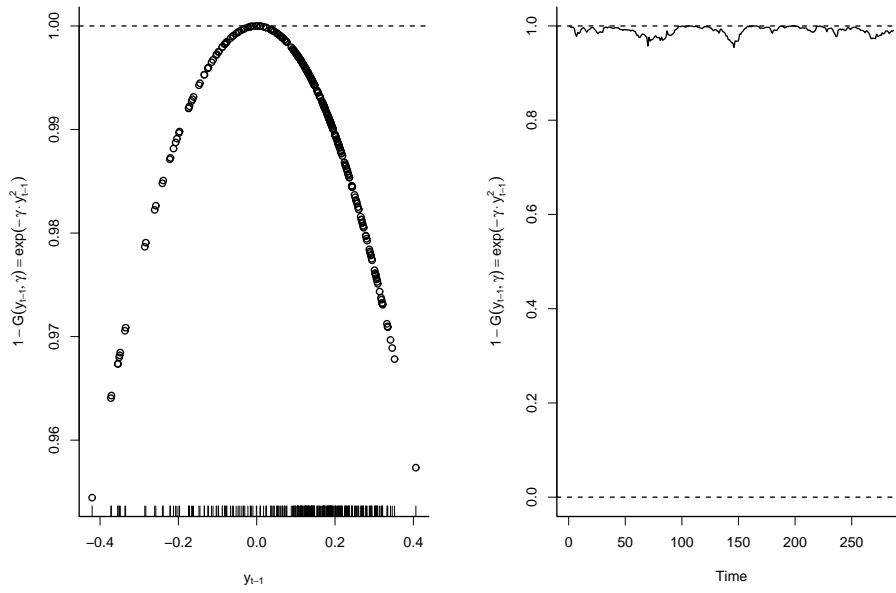


Figure 6: Left panel: One minus transition function against transition variable. Right panel: One minus transition function against time.

The figure supports the results that, albeit the parameter estimate for γ leads to a reasonable looking transition function plotting it against time, $\hat{\gamma} = 0.264$ actually produces a random walk model and by this contradicts PPP caused by a degenerated transition function. Producing the same plots for T-STAR we obtain Figure 7.

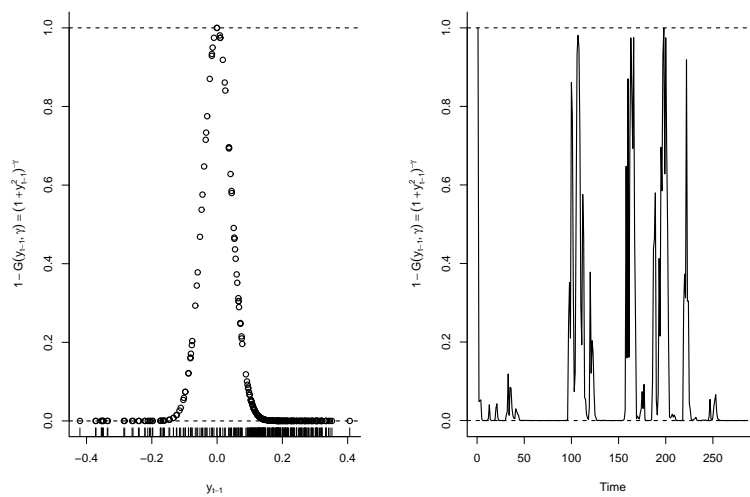


Figure 7: Left panel: One minus transition function against transition variable. Right panel: One minus transition function against time.

The estimation of a large $\hat{\gamma} = 275.284$ still produces a transition function that is by no means close to the limit for $\gamma \rightarrow \infty$ (see the properties in Section 2). In addition we see from the left panel that the estimated process is far more often in the stationary regime and becomes a random walk only on few occasions. This supports the PPP as a long run concept with a non-linear convergence towards an equilibrium. The estimation of the T-STAR model thus yields an advantage over ESTAR in terms of parameter interpretation.

It is noteworthy that the plot in Figure 6 is not unique for this particular data set but a common finding in empirically estimated ESTAR models. Kruse (2009) for example support these findings for other real exchange rates (see their Figure 1).

The panels in Figure 6 and Figure 7 also support the theoretical results that the estimation of γ heavily depends on the error term variance σ_ε derived in Section 2. Looking at the estimated values for γ in Table 13 and the left panels in Figures 6 and 7 we obtain a reasonable form of the transition function from a mathematical point of view. However plotting the transition function against time we see that the estimated function does not support the hypothesis that the data comes from the assumed data generating process. This supports that the estimation of γ is heavily influenced by the small error standard deviation. The estimated γ for the T-STAR model looks awkward at first. However, looking at the plots in Figure 7 this yields a transition function that is not degenerated in the sense that the actual range exploits its whole domain and it does not behave as in the limiting case (see Section 2). This supports the assumed data generating process, i.e. a globally stationary T-STAR model. The estimation of γ in the T-STAR case is by far not so heavily influenced from the small error standard deviation and thus we can extend the range of possible γ values for which we obtain a non-degenerated transition function. This also supports the conclusion that we can largely reduce the influence of σ_ε on γ by reformulating the transition function.

5 Conclusions

We have studied the ESTAR and T-STAR model, two competing models of the STAR family sharing the same characteristic properties of their transition functions. Due to their nonlinear structure, unidentified parameters occur for certain combinations of γ and σ^2 , the transition parameter and the error term variance, respectively. This phenomena has not been studied systematically before although it is of importance in applications.

In the ESTAR setting, very small values of σ^2 , among others, yield in particular an unidentified γ , making a consistent estimation of γ nearly impossible. In Theorem 2.4 we verified this by showing that the variance of the conditional Maximum Likelihood estimator $\hat{\gamma}$ tends to infinity as σ^2 vanishes. Hence, in order to estimate γ , somewhat unpleasant modifications need to be incorporate into the optimization routine.

In order to avoid this, we define the T-STAR model where γ becomes unidentified much later as $\sigma^2 \rightarrow 0$ compared to the ESTAR model. As a consequence, γ can be included in the parameter vector that is to be estimated. By deriving a linearity and a unit root test for the T-STAR model we support our opinion that this new model is indeed a worthy alternative, applicable to the same situations and should therefore be preferred to the ESTAR model.

This conclusion is illustrated by fitting both models to the same data set containing real exchange rates. The estimators one obtains in the ESTAR setting do not allow for a meaningful interpretation of the fitted model as one regime is basically switched off. One can clearly see that this is the result of the identification problem caused by a small error term variance. Contrary to that, the fitted

T-STAR models allows for switching between the two regimes, leading to a better fit, although the estimators for γ and σ^2 look not very promising without interpreting them in the right context.

As this text deals with a topic that is not well studied yet, there are many possible open questions and possibilities how to go from here. The most interesting question at the moment is whether it is possible to quantify and compare the regions for which (γ, σ) cause the identification problem in both regimes. A more theoretical, in depth study of the T-STAR model would be needed for this.

A Appendix: Proofs and Technical Lemmas

Proof of Lemma 2.3.

In order to show that $\mathbb{E} \left[\exp(-n\gamma y_t^2) y_t^{2k+1} \right] = 0$ for all $t \geq 0, k \in \mathbb{N}_0$ and $n \in \mathbb{N}_0$ we verify that $y_t^{2k+1} \exp(-n\gamma y_t^2)$ has a symmetric density for all $t \geq 0, k \in \mathbb{N}_0$ and $n \in \mathbb{N}_0$ as long as the density of ε_t is symmetric around zero for all t . This is done by first showing that y_t has a symmetric density for all $t \geq 0$, followed by an application of Lemma A.1 with $a = 0, b = 1$ and $c = n\gamma > 0$.

An inductive argument is used to prove that the density of y_t is symmetric around zero for all t . The initial variable $y_0 = \varepsilon_0$ has a symmetric density by choice. Now, let $t \geq 1$ and assume that y_{t-1} has a symmetric density and recall from (5) that

$$y_t = [\psi + \varphi \exp(-\gamma y_{t-1}^2)] y_{t-1} + \varepsilon_t, \quad t \geq 1.$$

We know from Lemma A.1 that $[\psi + \varphi \exp(-\gamma y_{t-1}^2)] y_{t-1}$ has a symmetric density around zero. As the same holds for ε_t by assumption, it follows from Lemma A.2 that also the density of y_t is symmetric around zero due to y_{t-1} and ε_t being independent.

In order to show that the left hand side of (7) behaves asymptotically like σ^{2k} as σ goes to infinity, we use again an inductive argument. First note that

$$\left| \mathbb{E} \left(\exp(-n\gamma y_t^2) y_t^{2k} \right) \right| \leq \mathbb{E} \left(y_t^{2k} \right)$$

and recall that the moment generating function of a normally distributed random variable X is given by $M_X(t) = \exp(\mu t + q/2 * \sigma^2 t^2)$ with

$$\mathbb{E}(X^n) = M_X^{(n)}(0). \tag{20}$$

Let $k \geq 1$ and $t = 0$. As y_0 is normally distributed with mean zero (see (6)) we obtain from (20)

$$\mathbb{E}(y_0^{2k}) = c_k \sigma^{2k}$$

for a constant c_k , as $M_{y_0}^{2k}(0) = c_k \sigma^{2k}$.

Assume now that (7) holds for $t - 1$. Then

$$\mathbb{E} \left(y_t^{2k} \right) = \mathbb{E} \left((\psi + \varphi \exp(-\gamma y_{t-1}^2))^{2k} y_{t-1}^{2k} \right) + \mathbb{E}(\varepsilon_t^{2k})$$

as the expected value of any product of y_{t-1} and ε_t vanished due to independence. As

$$(\psi + \varphi \exp(-\gamma y_{t-1}^2))^{2k} \leq (|\Psi| + |\varphi| \exp(-\gamma y_{t-1}^2))^{2k} \leq 1$$

we obtain

$$\mathbb{E} \left(y_t^{2k} \right) \leq \mathbb{E} \left(y_{t-1}^{2k} \right) + \mathbb{E}(\varepsilon_t^{2k}) = c_k \sigma^{2k} + c_k \sigma^{2k}$$

due to normality of ε_t , which shows that (7) holds. \square

Lemma A.1.

Let X be a real valued random variable with symmetric density around zero. Then the density of $[a + b \exp(-cX^2)]X^{2k+1}$ for some $k \in \mathbb{N}_0$ and $a, b, c \in \mathbb{R}$, $ab \neq 0, c > 0$, is symmetric around zero, too.

Proof.

The result is obtained by applying theorems deriving the density of a transformed random variable (see e.g. Theorems 22.2 and 22.3 in Behnen and Neuhaus (1995)).

First let $a \in \mathbb{R} \setminus \{0\}$, $c > 0$ and $b = 0$ and define $g_k : \mathbb{R} \rightarrow \mathbb{R}$, $g_k(x) = ax^{2k+1}$ for some $k \in \mathbb{N}_0$ as well as $Y = g_k(X)$. Since $g'_k(x) = (2k+1)ax^{2k}$, $x \in \mathbb{R}$, and $g_k^{-1}(x) = (x/a)^{1/(2k+1)}$, $x \in \mathbb{R}$, we obtain

$$f_Y(y) = \begin{cases} 0 & \text{for } y = 0, \\ \frac{f_X(g_k^{-1}(y))}{|g'_k(g_k^{-1}(y))|} = \frac{f_X\left(y^{\frac{1}{2k+1}} a^{-\frac{1}{2k+1}}\right)}{(2k+1)\left(y^{\frac{1}{2k+1}} a^{-\frac{1}{2k+1}}\right)^{2k}} & \text{for } y \neq 0, \end{cases}$$

where f_X and f_Y denote the densities of X and Y , respectively. Note that $(-y/a)^{1/(2k+1)} = (-1)^{1/(2k+1)}(y/a)^{1/(2k+1)} = -(y/a)^{1/(2k+1)}$ which implies, for $y \neq 0$,

$$f_Y(-y) = \frac{f_X\left(-y^{\frac{1}{2k+1}} a^{-\frac{1}{2k+1}}\right)}{(2k+1)\left(-y^{\frac{1}{2k+1}} a^{-\frac{1}{2k+1}}\right)^{2k}} = \frac{f_X\left(y^{\frac{1}{2k+1}} a^{-\frac{1}{2k+1}}\right)}{(2k+1)\left(y^{\frac{1}{2k+1}} a^{-\frac{1}{2k+1}}\right)^{2k}} = f_Y(y), \quad (21)$$

where the second last equality is due to the symmetry of f_X .

Now consider the case $b \neq 0$ which is incomparable more complex. Let $a \in \mathbb{R}$ and $c > 0$ with $g_k : \mathbb{R} \rightarrow \mathbb{R}$, $g_k(x) = (a + b \exp(-cx^2))x^{2k+1}$ for some $k \in \mathbb{N}_0$ and $Y = g_k(X)$. Contrary to the case $b = 0$, g_k can now change its monotonic behavior and might therefore be only piecewise invertible.

We know that, for all $k \in \mathbb{N}_0$,

- (i) g_k is continuous on \mathbb{R} with $g_k(0) = 0$,
- (ii) g_k is point symmetric around zero,
- (iii) by the properties of the exponential function,

$$\lim_{x \rightarrow \infty} g_k(x) = \lim_{x \rightarrow \infty} ax^{2k+1} + b \frac{x^{2k+1}}{e^{cx^2}} = \lim_{x \rightarrow \infty} ax^{2k+1} = \begin{cases} -\infty & \text{for } a < 0, \\ 0 & \text{for } a = 0, \\ \infty & \text{for } a > 0. \end{cases} \quad (22)$$

- (iv) the monotonic behavior can be summarized in the following table where $\xi_k = \frac{2}{2k+1}e^{-\frac{2k+3}{2}}$:

| | | |
|---------|---|---|
| $b > 0$ | $a > \xi_k b > 0$ | g_k is strictly monotone increasing |
| $b > 0$ | $a < -b < 0$ | g_k is strictly monotone decreasing |
| $b > 0$ | $a \in [-b, \xi_k b]$ $= [-b, 0]$ $\cup (0, \xi_k b]$ | g_k changes its monotone behavior twice, starting with being strictly decreasing g_k changes its monotone behavior four times, starting with being strictly increasing |
| $b < 0$ | $a > -b > 0$ | g_k is strictly monotone increasing |
| $b < 0$ | $a < \xi_k b < 0$ | g_k is strictly monotone decreasing |
| $b < 0$ | $a \in [\xi_k b, -b]$ $= [\xi_k b, 0)$ $\cup [0, -b]$ | g_k changes its monotone behavior four times, starting with being strictly decreasing g_k changes its monotone behavior twice, starting with being strictly increasing |

If g_k changes its monotonic behavior twice, it always happens at

$$w_{1,2} = \pm \sqrt{\frac{-W_0\left(-\frac{a}{b} \frac{2k+1}{2} e^{\frac{2k+1}{2}}\right) + \frac{2k+1}{2}}{c}}$$

no matter which parameter combination for a and b we consider. Here, W_0 denotes the principal branch of the Lambertsche W-function with domain $[-\exp(-1), \infty)$, i.e. the function that satisfies $x = W_0(x) \exp(W_0(x))$. If g_k changes its behavior four times, it additionally happens at

$$w_{3,4} = \pm \sqrt{\frac{-W_{-1}\left(-\frac{a}{b} \frac{2k+1}{2} e^{\frac{2k+1}{2}}\right) + \frac{2k+1}{2}}{c}}$$

where W_{-1} is the second real branch of the W-function defined on $[\exp(-1), 0)$.

Figures 8 and 9 illustrate the function g_0 for different parameters a and b with $c = 1$. While we choose $b = 1 > 0$, $a = -1.2 < -b$ in Figure 8, the latter corresponds to $b = 1 > 0$, $a = 0.2 \in (0, -\xi_0 b] = (0, 0.4463]$.

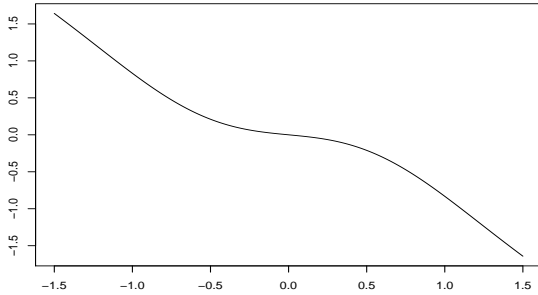


Figure 8: g_0 with $a = -1.2$ and $b = 1$.

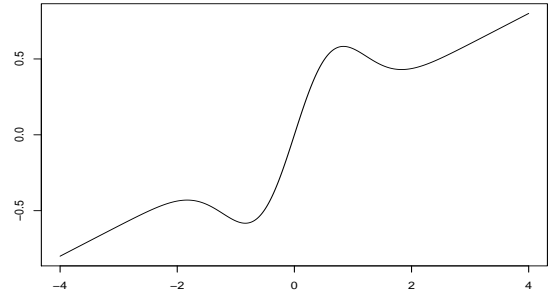


Figure 9: g_0 with $a = 0.2$ and $b = 1$.

Properties (i)-(iii) are obvious. We verify property (iv) by writing, for $x \in \mathbb{R}$,

$$\begin{aligned} g'_k(x) &= (2k+1)bx^{2k} \left[\frac{a}{b} + \exp(-cx^2) \left(1 - \frac{2}{2k+1} cx^2 \right) \right] \\ &= (2k+1)bx^{2k} \left[\frac{a}{b} - h_k(x) \right], \end{aligned}$$

with

$$h_k : \mathbb{R} \rightarrow \mathbb{R}, \quad h_k(x) = -\frac{2k+1}{2} \exp(-cx^2) \left(\frac{2k+1}{2} - cx^2 \right).$$

Let $k \geq 0$ and assume $b > 0$.

Then $(2k+1)bx^{2k} > 0$ for $x \in \mathbb{R} \setminus \{0\}$. Hence, $g'_k(x) > 0, x \in \mathbb{R} \setminus \{0\}$, (i.e. strictly monotone increasing) if $a/b > h(x)$ for all $x \in \mathbb{R} \setminus \{0\}$. As

$$\max_{x \in \mathbb{R}} h(x) = \max \left\{ h(0), h \left(\pm \sqrt{\frac{2k+3}{2c}} \right) \right\} = \frac{2}{2k+1} \exp \left(-\frac{2k+3}{2} \right) =: \xi_k,$$

we obtain a strictly increasing g_k for $a/b > \xi_k$ and by a similar argument a strictly decreasing g_k as long as $a/b < -1$ since

$$\min_{x \in \mathbb{R}} h(x) = h(0) = -1.$$

For $a/b \in [-1, \xi_k]$ the monotone behavior changes, driven by the sign of a (see (22)). Note that $g'_k(x) = 0$ whenever $x = 0$ ($k \geq 1$) or

$$\frac{a}{b} + \left(\frac{2k+1}{2} - cx^2 \right) \frac{2}{2k+1} \exp(-cx^2) = 0$$

which is equivalent to

$$\left(\frac{2k+1}{2} - cx^2 \right) \exp \left(\frac{2k+1}{2} - cx^2 \right) = -\frac{a}{b} \frac{2k+1}{2} \exp \left(\frac{2k+1}{2} \right). \quad (23)$$

For solving (23) we need to consider two different cases. Note that for $a \in (0, \xi_k b]$, the right hand side of (23) is contained in $[-\exp(-1), 0)$, hence in the range where W has two real-values branches, denoted by W_0 and W_{-1} . Therefore, from (23), for $j = 0, -1$,

$$\begin{aligned} \frac{2k+1}{2} - cx^2 &= W_j \left(-\frac{a}{b} \frac{2k+1}{2} \exp \left(\frac{2k+1}{2} \right) \right) \\ \Leftrightarrow w_{1,2,3,4} &= \pm \sqrt{\frac{-W_j \left(-\frac{a}{b} \frac{2k+1}{2} \exp \left(\frac{2k+1}{2} \right) \right) + \frac{2k+1}{2}}{c}} \end{aligned} \quad (24)$$

which are well defined as

$$-W_0 \left(-\frac{a}{b} \frac{2k+1}{2} \exp \left(\frac{2k+1}{2} \right) \right) \geq -W_0 \left(\frac{2k+1}{2} \exp \left(\frac{2k+1}{2} \right) \right) \geq -\frac{2k+1}{2}$$

and

$$-W_{-1} \left(-\frac{a}{b} \frac{2k+1}{2} \exp \left(\frac{2k+1}{2} \right) \right) \geq -W_{-1}(-\exp(-1)) = 1.$$

For $a \in [-1, 0]$ the argument of W_j in (24) is in the domain of only one real branch, namely W_0 , so that (23) has only two solution leading to two monotone changes of g_k .

An analogue argument shows the behavior of g_k if $b < 0$.

For those combinations of a and b where g_k is strictly monotone, symmetry of f_Y can be derived in the same way as in (21). Note that g'_k is symmetric around zero and that g_k^{-1} is point symmetric around zero as it shares this property with g . Thus, for $y \in \mathbb{R}$,

$$f_Y(-y) = \frac{f_X(g_k^{-1}(-y))}{|g'_k(g_k^{-1}(-y))|} = \frac{f_X(-g_k^{-1}(-y))}{|g'_k(-g_k^{-1}(-y))|} = \frac{f_X(g_k^{-1}(y))}{|g'_k(g_k^{-1}(y))|} = f_Y(y).$$

For the remaining parameters combinations of a and b we present as an example the argument for $b > 0$ and $a \in (0, \xi_k b]$. Define

$$G_1 = (-\infty, -z_2) \cup (z_2, \infty), \quad G_2 = (-z_2, -z_1) \cup (z_1, z_2) \quad \text{and} \quad G_3 = (-z_1, z_1)$$

where $z_1 = w_1$ and $z_2 = w_3$. On the open sets G_i , $i = 1, \dots, 3$, $g_{k,i} = g_k \mathbb{1}_{G_i}$, $i = 1, \dots, 3$ is strictly monotone with derivatives and inverse functions $g'_{k,i}$ and $g_{k,i}^{-1}$, respectively. Hence, from Theorem 22.3 of Behnen and Neuhaus (1995),

$$f_Y(y) = \sum_{i=1}^3 \frac{f(g_{k,i}^{-1}(y))}{|g'_{k,i}(g_{k,i}^{-1}(y))|} \mathbb{1}_{y \in H_i}, \quad y \in \mathbb{R},$$

for $H_i = g_{k,i}(G_i)$. As $H_1 \cup H_2 \cup H_3 = \mathbb{R}$, as $g'_{k,i}$, $i = 1, \dots, 3$ are symmetric around zero and as $g_{k,i}^{-1}$, $i = 1, \dots, 3$ are point symmetric around zero, we obtain $f_Y(-y) = f_Y(y)$ for all $y \in \mathbb{R}$. \square

Lemma A.2.

Let X and Y be independent real valued random variables with densities f_X and f_Y , respectively. Symmetries of f_X around $c \in \mathbb{R}$ ($f_X(x+c) = f_X(-x+c)$, $x \in \mathbb{R}$) and f_Y around $d \in \mathbb{R}$ then imply that the convolution density f_Z of $Z = X + Y$ is symmetric around $c + d$.

Proof.

The density f_Z of Z is given by

$$f_Z(z) = \int_{\mathbb{R}} f_X(z-x)f_Y(x) dx, \quad z \in \mathbb{R}.$$

Hence, for all $x \in \mathbb{R}$,

$$\begin{aligned} f_Z(-x+c+d) &= \int_{\mathbb{R}} f_X(-x+c+d-y)f_Y(y) dy = \int_{\mathbb{R}} f_X(-(x-d+y)+c)f_Y(y) dy \\ &= \int_{\mathbb{R}} f_X(x-d+y+c)f_Y(y) dy = \int_{\mathbb{R}} f_X(x+c+d+y-2d)f_Y(y) dy \\ &= \int_{\mathbb{R}} f_X(x+c+d-z)f_Y(-z+2d) dz = \int_{\mathbb{R}} f_X(x+c+d-z)f_Y(-(z-d)+d) dz \\ &= \int_{\mathbb{R}} f_X(x+c+d-z)f_Y(z-d+d) dz = f_Z(x+c+d). \quad \square \end{aligned}$$

Proof of Theorem 2.4.

Let $\hat{\beta}_n$ be the conditional Maximum Likelihood estimator of $\beta = (\psi, \varphi, \gamma)$. Then by Theorem 3.2 of Tjøstheim (1986)

$$n^{1/2}(\hat{\beta}_n - \beta) \xrightarrow{d} N(0, \sigma^2 U^{-1})$$

and where the matrix U is given by

$$U = \begin{pmatrix} \mathbb{E}[y_{t-1}^2] & \mathbb{E}[y_{t-1}^2 \exp(-\gamma y_{t-1}^2)] & -2\mathbb{E}[y_{t-1}^3 \exp(-\gamma y_{t-1}^2)] \\ \mathbb{E}[y_{t-1}^2 \exp(-\gamma y_{t-1}^2)] & \mathbb{E}[y_{t-1}^2 \exp(-2\gamma y_{t-1}^2)] & -2\mathbb{E}[y_{t-1}^3 \exp(-2\gamma y_{t-1}^2)] \\ -2\mathbb{E}[y_{t-1}^3 \exp(-\gamma y_{t-1}^2)] & -2\mathbb{E}[y_{t-1}^3 \exp(-2\gamma y_{t-1}^2)] & 4\mathbb{E}[y_{t-1}^4 \exp(-\gamma y_{t-1}^2)] \end{pmatrix}.$$

In order to obtain the limiting behavior of $\text{Var}(\hat{\gamma})$ we therefore study

$$\sigma^2 (U^{-1})_{33} = \frac{\sigma^2}{\det(U)} \det \begin{pmatrix} u_{11} & u_{12} \\ u_{21} & u_{22} \end{pmatrix}. \quad (25)$$

Because of Lemma A.1, which states that $\mathbb{E} \left[[\psi + \varphi \exp(-\gamma y_{t-1}^2)]^{2k+1} y_{t-1}^{2k+1} \right] = 0$, the matrix U boils down in our situation to

$$U = \begin{pmatrix} \mathbb{E}[y_{t-1}^2] & \mathbb{E}[y_{t-1}^2 \exp(-\gamma y_{t-1}^2)] & 0 \\ \mathbb{E}[y_{t-1}^2 \exp(-\gamma y_{t-1}^2)] & \mathbb{E}[y_{t-1}^2 \exp(-2\gamma y_{t-1}^2)] & 0 \\ 0 & 0 & 4\mathbb{E}[y_{t-1}^4 \exp(-\gamma y_{t-1}^2)] \end{pmatrix}$$

so that

$$\det(U) = u_{33} \det \begin{pmatrix} u_{11} & u_{12} \\ u_{21} & u_{22} \end{pmatrix}$$

and hence

$$\sigma^2 (U^{-1})_{33} = \sigma^2 u_{33}^{-1} = \frac{\sigma^2}{4\mathbb{E}[y_{t-1}^4 \exp(-\gamma y_{t-1}^2)]}.$$

Applying (7) to the denominator we get

$$\sigma^2 (U^{-1})_{33} \sim \frac{c\sigma^2}{\sigma^4} = \frac{1}{\sigma^2} \rightarrow \infty, \text{ for } \sigma \downarrow 0,$$

which finishes the proof. \square

Proof of Theorem 3.2.

We consider the asymptotic behavior of the least squares estimator $\hat{\beta} = (\hat{\delta}_1^2, \hat{\delta}_1^4, \hat{\delta}_1^6)$. Since under the null $\Delta y_t = u_t$ the OLS estimator $\hat{\beta}$ can be written as

$$\hat{\beta} = \underbrace{\left(\sum_{t=1}^T x_t' x_t \right)^{-1}}_I \underbrace{\sum_{t=1}^T x_t' u_t}_II.$$

The elements in this matrix equation read for part I

$$\sum_{t=1}^T x_t' x_t = \begin{bmatrix} \sum_{t=1}^T y_{t-1}^6 & \sum_{t=1}^T y_{t-1}^8 & \sum_{t=1}^T y_{t-1}^{10} \\ \sum_{t=1}^T y_{t-1}^8 & \sum_{t=1}^T y_{t-1}^{10} & \sum_{t=1}^T y_{t-1}^{12} \\ \sum_{t=1}^T y_{t-1}^{10} & \sum_{t=1}^T y_{t-1}^{12} & \sum_{t=1}^T y_{t-1}^{14} \end{bmatrix} \quad (26)$$

and for part II

$$\sum_{t=1}^T x'_t u_t = \begin{bmatrix} \sum_{t=1}^T y_{t-1}^3 u_t & \sum_{t=1}^T y_{t-1}^5 u_t & \sum_{t=1}^T y_{t-1}^7 u_t \end{bmatrix}. \quad (27)$$

In order to determine the asymptotic behavior we have to scale the OLS estimator $\hat{\beta}$ properly. We thus multiply $\hat{\beta}$ with the scaling matrix $\Gamma = \text{diag}(T^2, T^3, T^4)$ and obtain

$$\left[\Gamma^{-1} \sum_{t=1}^T x'_t x_t \Gamma^{-1} \right]^{-1} \left[\Gamma^{-1} \sum_{t=1}^T x'_t u_t \right] = \Gamma \hat{\beta}.$$

Now the asymptotic behavior of the first part of $\Gamma \hat{\beta}$ follows directly from Lemma A.3 and the behavior of the second part follows from the convergence to stochastic integrals for products of I(1) variables (Theorem 4.2 from Hansen (1992) the CMT and Itô's Lemma). This yields the following general result for $i \in \mathbb{N}_{>0}$

$$\frac{1}{T^{(i+1)/2}} \sum_{t=1}^T y_{t-1}^i u_t \Rightarrow \int_0^1 B^i(r) dB(r) = \sigma^{(i+1)} \left\{ \frac{1}{(i+1)} \int_0^1 B(1)^{(i+1)} - \frac{i}{2} \int_0^1 B(r) dr \right\}.$$

Given these results the OLS estimator converges as $T \rightarrow \infty$ as follows

$$\Gamma \hat{\beta} = \left[\Gamma^{-1} \sum_{t=1}^T x'_t x_t \Gamma^{-1} \right]^{-1} \left[\Gamma^{-1} \sum_{t=1}^T x'_t u_t \right] \Rightarrow \mathbf{Q}^{-1} \mathbf{v}$$

where

$$\mathbf{Q} = \begin{bmatrix} \sigma^6 \int_0^1 B^6(r) dr & \sigma^8 \int_0^1 B^8(r) dr & \sigma^{10} \int_0^1 B^{10}(r) dr \\ 0 & 0 & 0 \\ \sigma^8 \int_0^1 B^8(r) dr & \sigma^{10} \int_0^1 B^{10}(r) dr & \sigma^{12} \int_0^1 B^{12}(r) dr \\ 0 & 0 & 0 \\ \sigma^{10} \int_0^1 B^{10}(r) dr & \sigma^{12} \int_0^1 B^{12}(r) dr & \sigma^{14} \int_0^1 B^{14}(r) dr \\ 0 & 0 & 0 \end{bmatrix}$$

and

$$\mathbf{v} = \left[\sigma^4 \left\{ \frac{1}{4} \int_0^1 B(1)^4 - \frac{3}{2} \int_0^1 B(r) dr \right\} \quad \sigma^6 \left\{ \frac{1}{6} \int_0^1 B(1)^6 - \frac{5}{2} \int_0^1 B(r) dr \right\} \quad \sigma^8 \left\{ \frac{1}{8} \int_0^1 B(1)^8 - \frac{7}{2} \int_0^1 B(r) dr \right\} \right]'$$

The scaled F -statistic we are concerned with reads

$$F^* = \frac{1}{3} \Gamma \hat{\beta}' [\hat{\sigma}^2 \Gamma (X' X)^{-1} \Gamma]^{-1} \Gamma \hat{\beta}.$$

By the law of large numbers it is easy to show that under the null as $T \rightarrow \infty$

$$\hat{\sigma}^2 = \frac{1}{T-4} \sum_{t=1}^T \left(\Delta y_t - \hat{\delta}_0 y_{t-1}^3 - \hat{\delta}_1 y_{t-1}^5 - \hat{\delta}_2 y_{t-1}^7 \right)^2 \xrightarrow{P} \sigma^2.$$

The test statistic F^* has the limiting distribution

$$F^* \Rightarrow \frac{1}{3\sigma^2} (\mathbf{Q}^{-1} \mathbf{v})' (\mathbf{Q}^{-1})^{-1} (\mathbf{Q}^{-1} \mathbf{v}) = \frac{1}{3\sigma^2} \mathbf{v}' \mathbf{Q}^{-1} \mathbf{v}$$

as stated in the theorem.

Under the alternative Δy_t and $y_{t-1}^i, \forall i \in \mathbb{N}_{>0}$ are $I(0)$ and thus it is readily seen that

$$\frac{1}{T} \sum_{t=1}^T \Delta y_t = \mathcal{O}_P(1); \quad \frac{1}{T} \sum_{t=1}^T y_{t-1}^i = \mathcal{O}_P(1);$$

are bounded in probability. Furthermore the innovation process u_t is by assumption $I(0)$ and thus

$$\frac{1}{T} \sum_{t=1}^T u_t = \mathcal{O}_P(1)$$

as well. Consider first the behavior of the OLS estimate

$$\hat{\beta} = \left(\underbrace{\sum_{t=1}^T x_t' x_t}_I \right)^{-1} \underbrace{\sum_{t=1}^T x_t' u_t}_{II}.$$

The part I reads as in (26)

$$T \frac{1}{T} \sum_{t=1}^T x_t' x_t = T \mathcal{O}_P(1) = \mathcal{O}_P(T).$$

The part II can be written as

$$\sum_{t=1}^T x_t' u_t = \sum_{t=1}^T x_t' \sum_{t=1}^T u_t = T \frac{1}{T} \left(\sum_{t=1}^T x_t' \sum_{t=1}^T u_t \right) = T \underbrace{\frac{1}{T} \sum_{t=1}^T x_t'}_{\mathcal{O}_P(1)} T \underbrace{\frac{1}{T} \sum_{t=1}^T u_t}_{\mathcal{O}_P(1)} = T^2 \mathcal{O}_P(1) = \mathcal{O}_P(T^2).$$

For the OLS estimate $\hat{\beta}$ follows that

$$(\mathcal{O}_P(T))^{-1} \mathcal{O}_P(T^2) = (T \mathcal{O}_P(1))^{-1} T^2 \mathcal{O}_P(1) = \frac{1}{T} T^2 \mathcal{O}_P(1) = T \mathcal{O}_P(1) = \mathcal{O}_P(T).$$

Analogously for the test statistic F^* it follows that

$$F^* = \frac{1}{3\hat{\sigma}^2} \hat{\beta}' (X'X) \hat{\beta} = \frac{1}{3\hat{\sigma}^2} \mathcal{O}_P(T) \mathcal{O}_P(T) \mathcal{O}_P(T) = \frac{1}{3\hat{\sigma}^2} \mathcal{O}_P(T^3).$$

Hence as $T \rightarrow \infty$ the test statistic F^* diverges to infinity. \square

Lemma A.3.

Let $y_t = y_{t-1} + \varepsilon_t$ be a random walk and let ε_t be as in Assumption 2.1 such that a functional central limit theorem (FCLT) applies. Then, for all $i \in \mathbb{N}_{>0}$,

$$\frac{1}{T^{(i+2)/2}} \sum_{t=1}^T y_{t-1}^i \Rightarrow \sigma^i \int_0^1 B^i(r) dr.$$

Proof.

Denote the partial sum process by $X_T(r) = \sum_{t=1}^{\lfloor Tr \rfloor} y_{t-1}$ with $r \in [0, 1]$ and $\lfloor \cdot \rfloor$ being the integer part of its argument. Moreover, let $S_T(r) = \sqrt{T}X_T(r)$. The FCLT then implies

$$S_T(r) \Rightarrow \sigma \int_0^1 B(r)dr, T \rightarrow \infty,$$

so that, for $i \in \mathbb{N}_{>0}$,

$$S_T^i(r) \Rightarrow \sigma^i \int_0^1 B^i(r)dr$$

by the continuous mapping theorem. Write $S_T^i(r) = T^{i/2}X_T^i(r)$ with $X_T^i(r) = T^{-i} \sum_{t=1}^{\lfloor Tr \rfloor} y_{t-1}^i$.

Then $S_T^i(r)$ reads

$$S_T^i(r) = \begin{cases} 0 & \text{for } 0 \leq r < 1/T \\ T^{(i/2-i)}y_1^i & \text{for } 1/T \leq r < 2/T \\ T^{(i/2-i)}y_2^i & \text{for } 2/T \leq r < 3/T \\ \vdots & \\ T^{(i/2-i)}y_T^i & \text{for } r = 1. \end{cases}$$

Therefore

$$\int_0^1 S_T^i(r) = T^{-1} \left[T^{(i/2-i)} \sum_{t=1}^T y_{t-1}^i \right] = T^{(-i/2-1)} \sum_{t=1}^T y_{t-1}^i = \frac{1}{T^{(i+2)/2}} \sum_{t=1}^T y_{t-1}^i.$$

If $T \rightarrow \infty$ it follows by the FCLT and the CMT

$$\frac{1}{T^{(i+2)/2}} \sum_{t=1}^T y_{t-1}^i \Rightarrow \sigma^i \int_0^1 B^i(r)dr. \quad \square$$

Proof of Theorem 3.3.

We have to show that the inner product of the regressor matrix including the additional regressors is asymptotically block diagonal (see e.g. [Hamilton \(1994\)](#) or [Hatanaka \(1996\)](#)). The inner product

$(X'X)$ of the regressor matrix from (19) reads

$$\begin{bmatrix} \sum_{t=1}^T y_{t-1}^6 & \sum_{t=1}^T y_{t-1}^8 & \sum_{t=1}^T y_{t-1}^{10} & \sum_{t=1}^T y_{t-1}^3 \Delta y_{t-1} & \cdots & \sum_{t=1}^T y_{t-1}^3 \Delta y_{t-p} \\ \sum_{t=1}^T y_{t-1}^8 & \sum_{t=1}^T y_{t-1}^{10} & \sum_{t=1}^T y_{t-1}^{12} & \sum_{t=1}^T y_{t-1}^5 \Delta y_{t-1} & \cdots & \sum_{t=1}^T y_{t-1}^5 \Delta y_{t-p} \\ \sum_{t=1}^T y_{t-1}^{10} & \sum_{t=1}^T y_{t-1}^{12} & \sum_{t=1}^T y_{t-1}^{14} & \sum_{t=1}^T y_{t-1}^7 \Delta y_{t-1} & \cdots & \sum_{t=1}^T y_{t-1}^7 \Delta y_{t-p} \\ \cdots & \cdots & \cdots & \cdots & \cdots & \cdots \\ \sum_{t=1}^T \Delta y_{t-1} y_{t-1}^3 & \sum_{t=1}^T \Delta y_{t-1} y_{t-1}^5 & \sum_{t=1}^T \Delta y_{t-1} y_{t-1}^7 & \sum_{t=1}^T \Delta y_{t-1}^2 & \cdots & \sum_{t=1}^T \Delta y_{t-1} \Delta y_{t-p} \\ \vdots & \vdots & \vdots & \vdots & \cdots & \vdots \\ \sum_{t=1}^T \Delta y_{t-p} y_{t-1}^3 & \sum_{t=1}^T \Delta y_{t-p} y_{t-1}^5 & \sum_{t=1}^T \Delta y_{t-p} y_{t-1}^7 & \sum_{t=1}^T \Delta y_{t-p} \Delta y_{t-1} & \cdots & \sum_{t=1}^T \Delta y_{t-p}^2 \end{bmatrix}.$$

Remember (see e.g. [Hamilton, 1994](#), p.517) that an $AR(p)$ process

$$(1 - \phi_1 L - \phi_2 L^2 - \dots - \phi_p L^p) y_t = \varepsilon_t$$

can be written equivalently as

$$\{(1 - \rho L) - (\zeta_1 L + \zeta_2 L^2 + \dots + \zeta_{p-1} L^{p-1})(1 - L)\} y_t = \varepsilon_t$$

where $\rho = \phi_1 + \phi_2 + \dots + \phi_p$ and $\zeta_j = -[\phi_{j+1} + \phi_{j+2} + \dots + \phi_p]$ for $j = 1, 2, \dots, p-1$. Under the assumption of a unit root, i.e. $\rho = 1$, the process can be written as

$$(\zeta_1 L - \zeta_2 L^2 - \dots - \zeta_{p-1} L^{p-1}) \Delta y_t = \varepsilon_t$$

or

$$\Delta y_t = e_t$$

where $e_t = (\zeta_1 L - \zeta_2 L^2 - \dots - \zeta_{p-1} L^{p-1})^{-1}$. The behavior of the process y_t is such that it fulfills proposition 17.3 in [Hamilton \(1994, p.505\)](#).

First, letting $e_t = y_t - y_{t-1}$ we obtain

$$\begin{bmatrix} \sum_{t=1}^T y_{t-1}^6 & \sum_{t=1}^T y_{t-1}^8 & \sum_{t=1}^T y_{t-1}^{10} & \sum_{t=1}^T y_{t-1}^3 e_{t-1} & \cdots & \sum_{t=1}^T y_{t-1}^3 e_{t-p} \\ \sum_{t=1}^T y_{t-1}^8 & \sum_{t=1}^T y_{t-1}^{10} & \sum_{t=1}^T y_{t-1}^{12} & \sum_{t=1}^T y_{t-1}^5 e_{t-1} & \cdots & \sum_{t=1}^T y_{t-1}^5 e_{t-p} \\ \sum_{t=1}^T y_{t-1}^{10} & \sum_{t=1}^T y_{t-1}^{12} & \sum_{t=1}^T y_{t-1}^{14} & \sum_{t=1}^T y_{t-1}^7 e_{t-1} & \cdots & \sum_{t=1}^T y_{t-1}^7 e_{t-p} \\ \cdots & \cdots & \cdots & \cdots & \cdots & \cdots \\ \sum_{t=1}^T e_{t-1} y_{t-1}^3 & \sum_{t=1}^T e_{t-1} y_{t-1}^5 & \sum_{t=1}^T e_{t-1} y_{t-1}^7 & \sum_{t=1}^T e_{t-1}^2 & \cdots & \sum_{t=1}^T e_{t-1} e_{t-p} \\ \vdots & \vdots & \vdots & \vdots & \cdots & \vdots \\ \sum_{t=1}^T e_{t-p} y_{t-1}^3 & \sum_{t=1}^T e_{t-p} y_{t-1}^5 & \sum_{t=1}^T e_{t-p} y_{t-1}^7 & \sum_{t=1}^T e_{t-p} e_{t-1} & \cdots & \sum_{t=1}^T e_{t-p}^2 \end{bmatrix}.$$

Using the results (c) and (e) stated in proposition 17.3 in [Hamilton \(1994, p.505\)](#) combined with the CMT and the results from theorem 3.2 we have

$$(X'X) \Rightarrow \begin{bmatrix} \mathbf{Q} & \mathbf{0} \\ \mathbf{0} & \mathbf{W} \end{bmatrix}$$

where

$$\mathbf{W} = \begin{bmatrix} \gamma_0 & \gamma_1 & \cdots & \gamma_{p-2} \\ \gamma_1 & \gamma_0 & \cdots & \gamma_{p-3} \\ \vdots & \vdots & \cdots & \vdots \\ \gamma_{p-2} & \gamma_{p-3} & \cdots & \gamma_0 \end{bmatrix}, \quad \gamma_j = E[(\Delta y_t)(\Delta y_{t-j})]$$

and where \mathbf{Q} is as given in Theorem 3.2 but with σ^k replaced by its long-run counterpart given by $\lambda = \sigma/(1 - \zeta_1 - \dots - \zeta_{p-1})$. Thus the inner product of the regressor matrix is asymptotically block diagonal and therefore the distribution of the coefficients δ_1^2 , δ_1^4 and δ_1^6 is independent of the distribution of the additional regressors.

Using similar arguments as in Theorem 3.2 it is straightforward to show that the test is consistent under (19). \square

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